

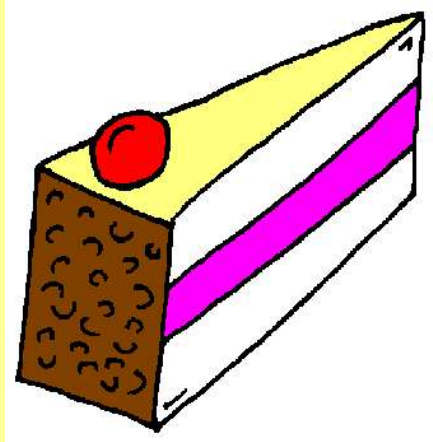


Reconnection in Space

Andris Valiulis

Uppsala, 2004-06-07

Docent lecture



after lecture - cake refreshments at 3rd floor

- ... include parts that my 1 year old daughter could understand
- ... include parts that nobody understands



research

Reconnection in space

- ✓ **Energy**
- ✓ **Magnetic field**
- ✓ **Reconnection concept**
- ✓ **Reconnection examples**
- ✓ **Reconnection research**

Energy

For those who want some proof that physicists are human, the proof is in the idiocy of all the different units which they use for measuring energy.

The Character of Physical Law (1967) R.P. Feynman.

J, °C, °F, K, degrees, eV, erg, cal, ...



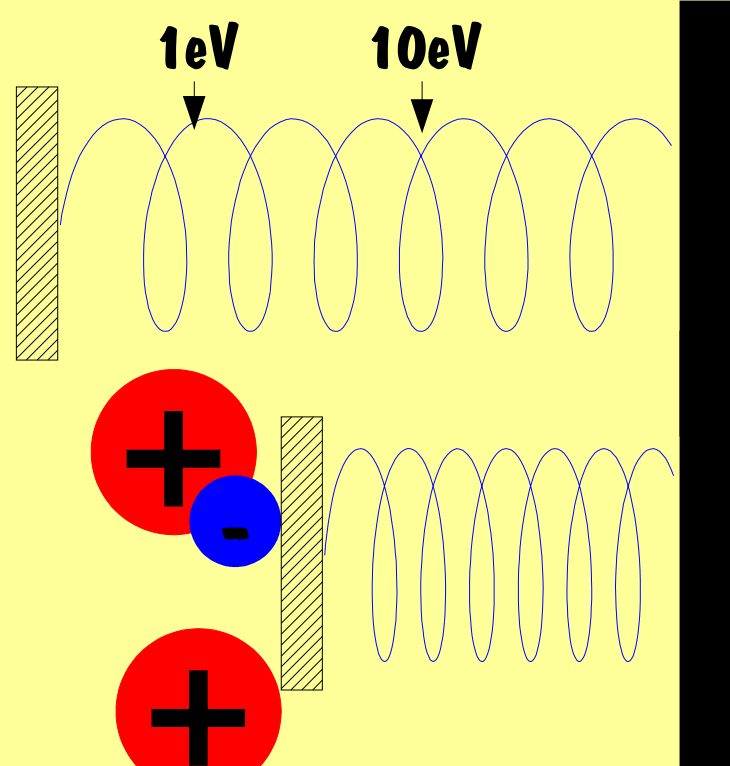
Measuring energy

atoms

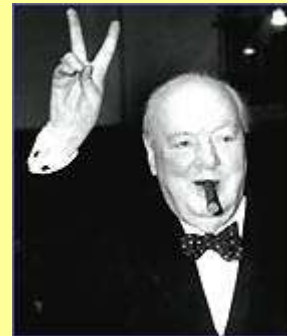
ions

electrons

photons



$h\nu$



1 eV – infrared

2 eV - visible light

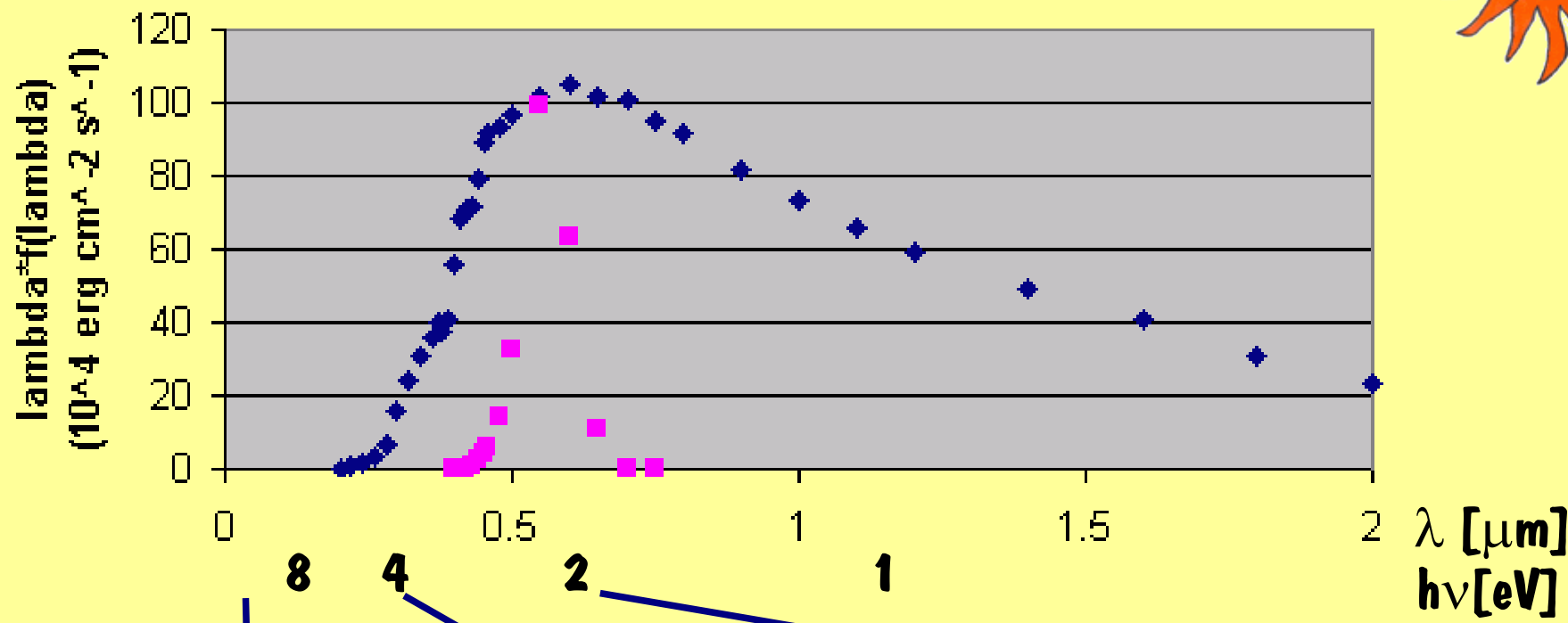
4 eV – ultraviolet

1keV – X rays

1MeV – gamma rays

Sun

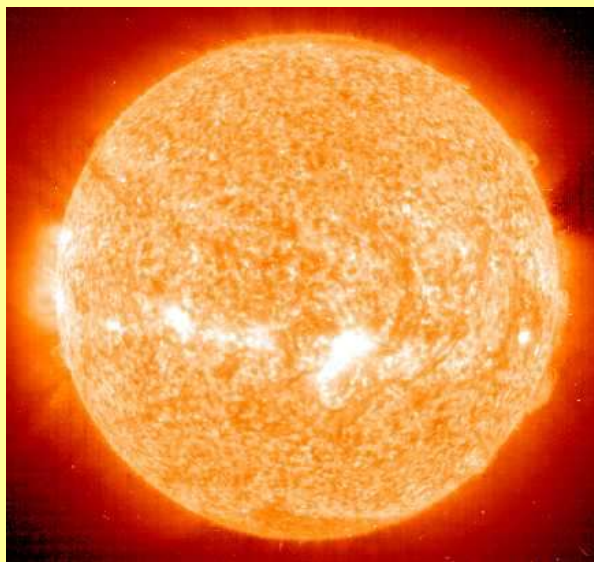
Surface temperature $\sim 6000\text{ }^\circ\text{C}$



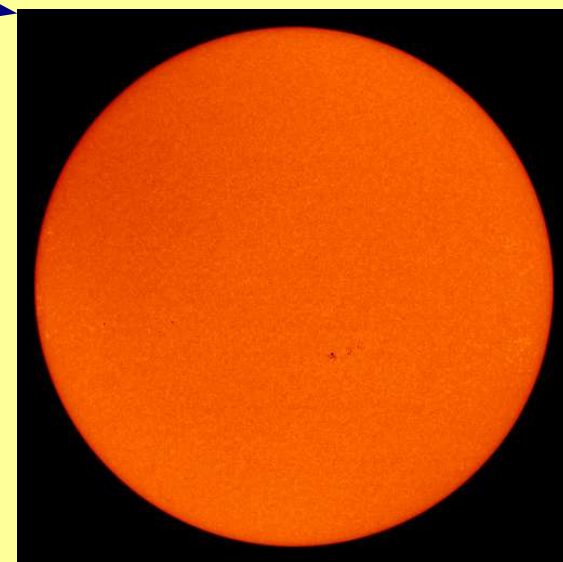
1 keV



4 eV

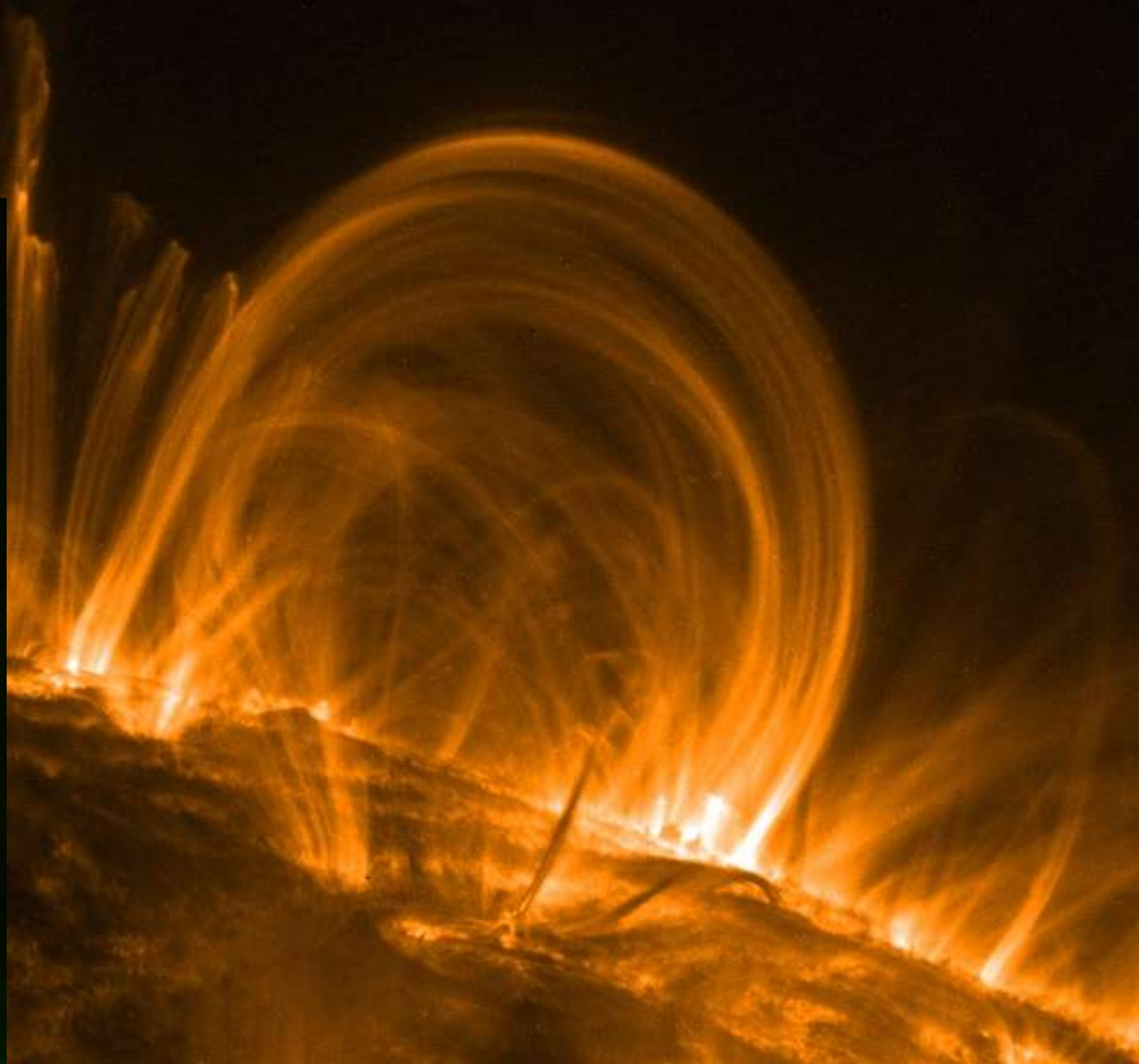
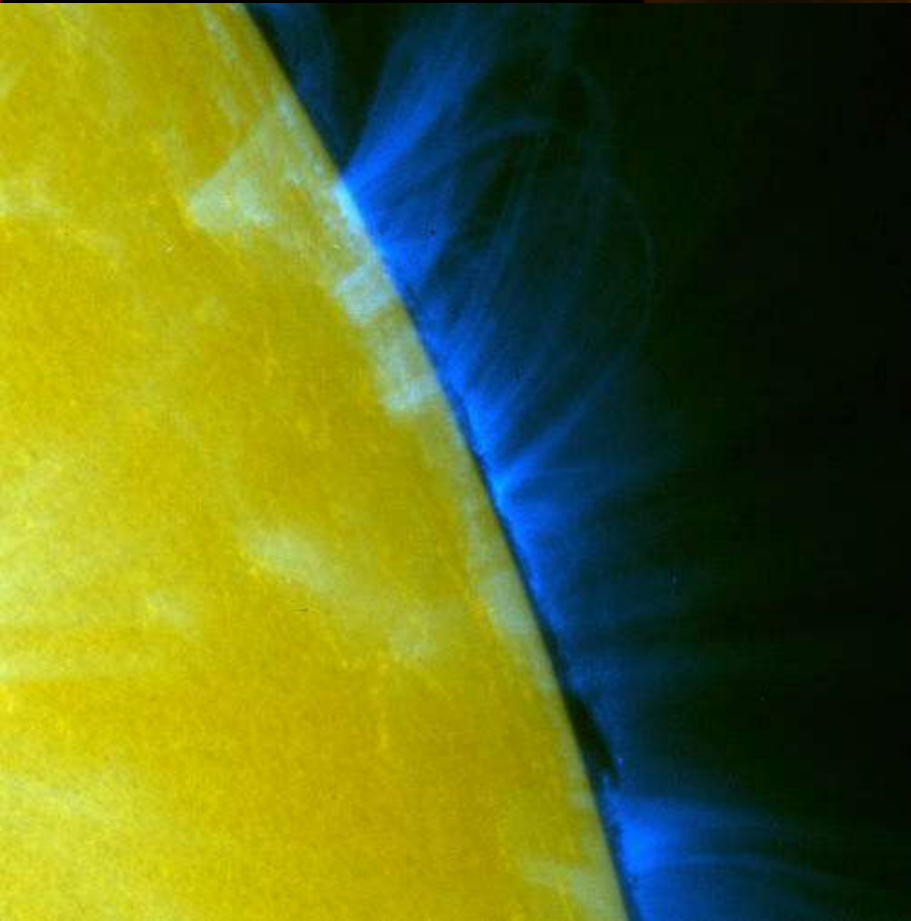


2 eV



**Highest energies in space above the Sun!
Energy should come from somewhere!**

Trace, 2eV, 7eV



Sunny day on Earth

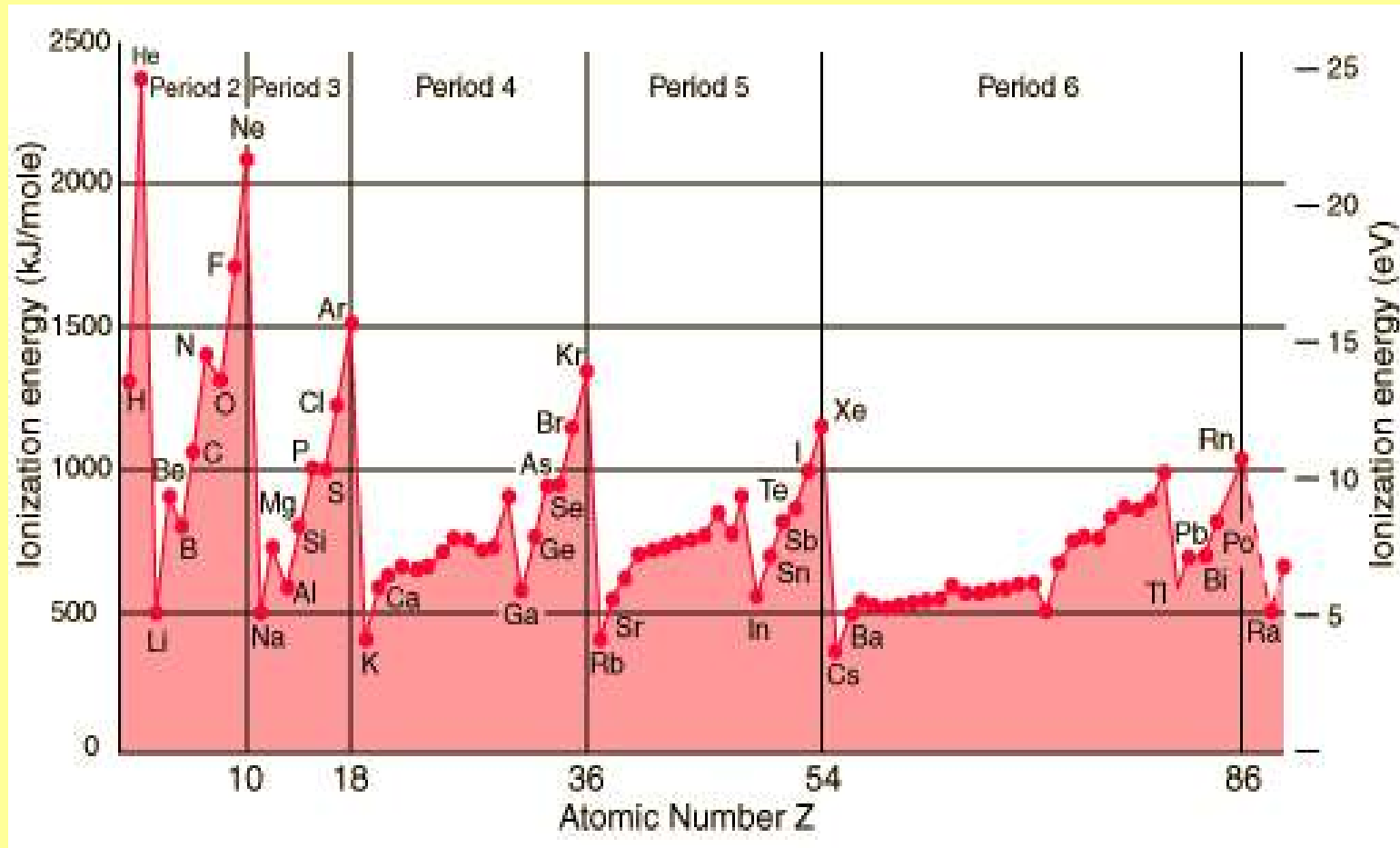
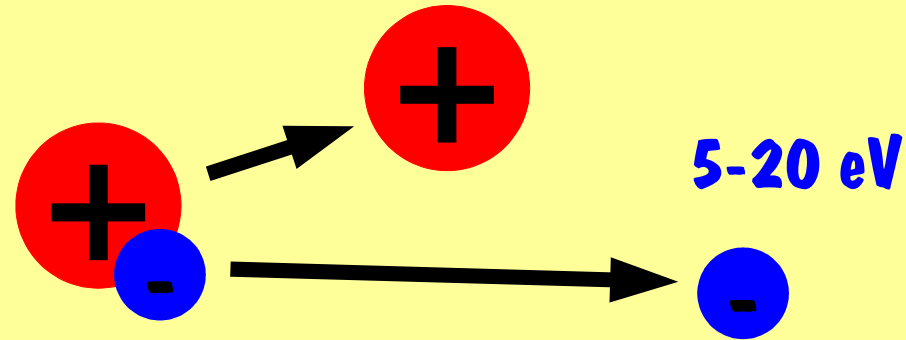


Particle energies

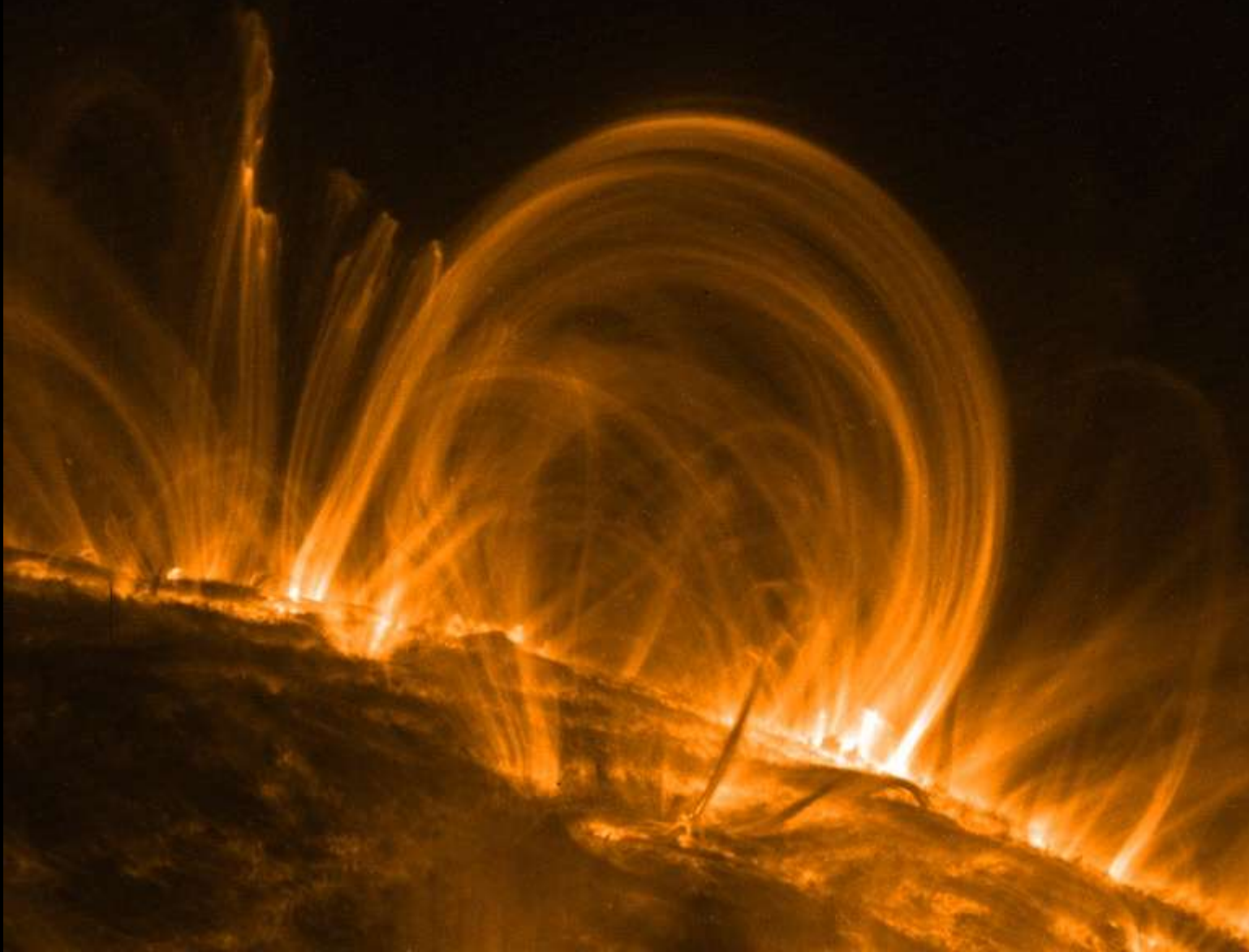
	$^{\circ}\text{C}$	eV
on Sun	6000	0.5
space around Sun	10^6	100
on Earth	30	0.026
space around Earth	10^3-10^8	0.1-10000
center of Sun	$15 \cdot 10^6$	1300

Ionization

Energy necessary to create an ion and electron from one atom

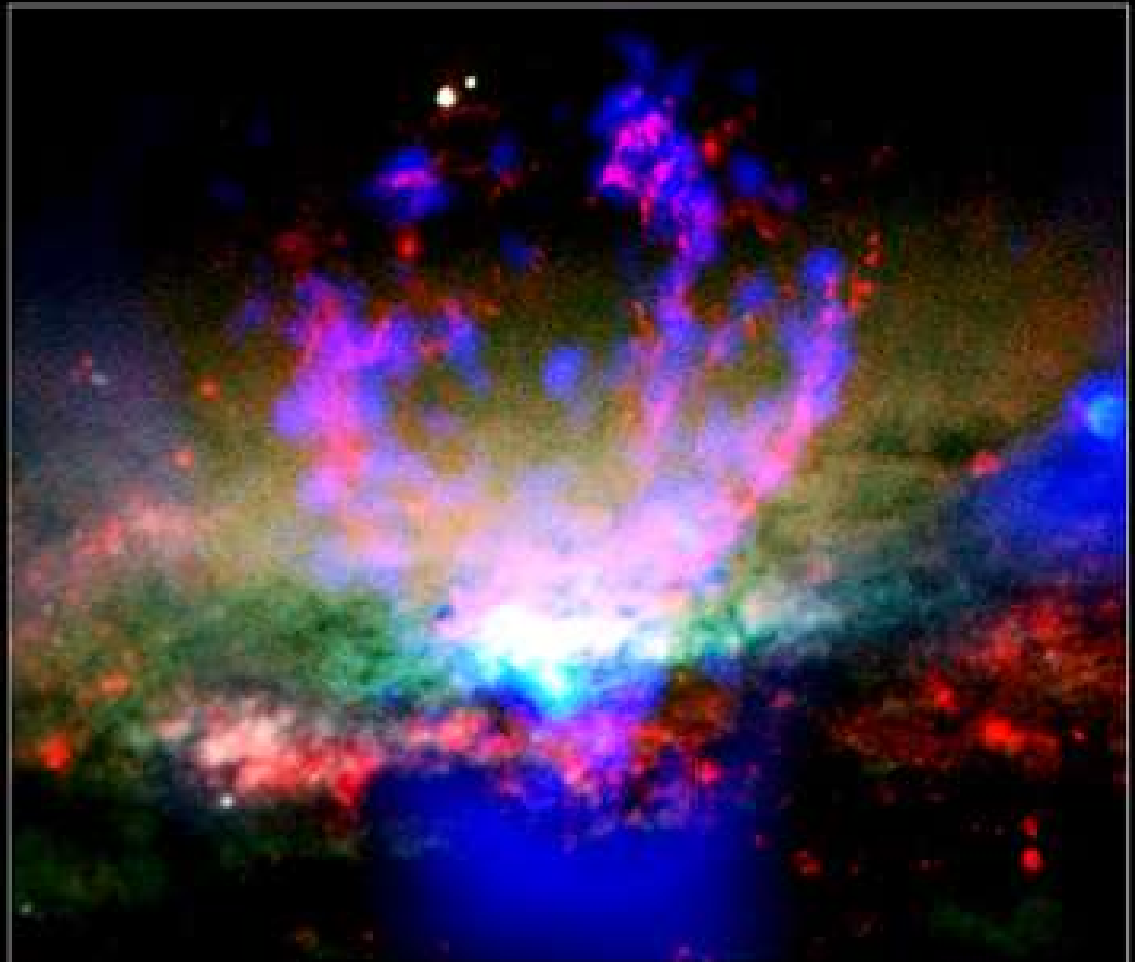
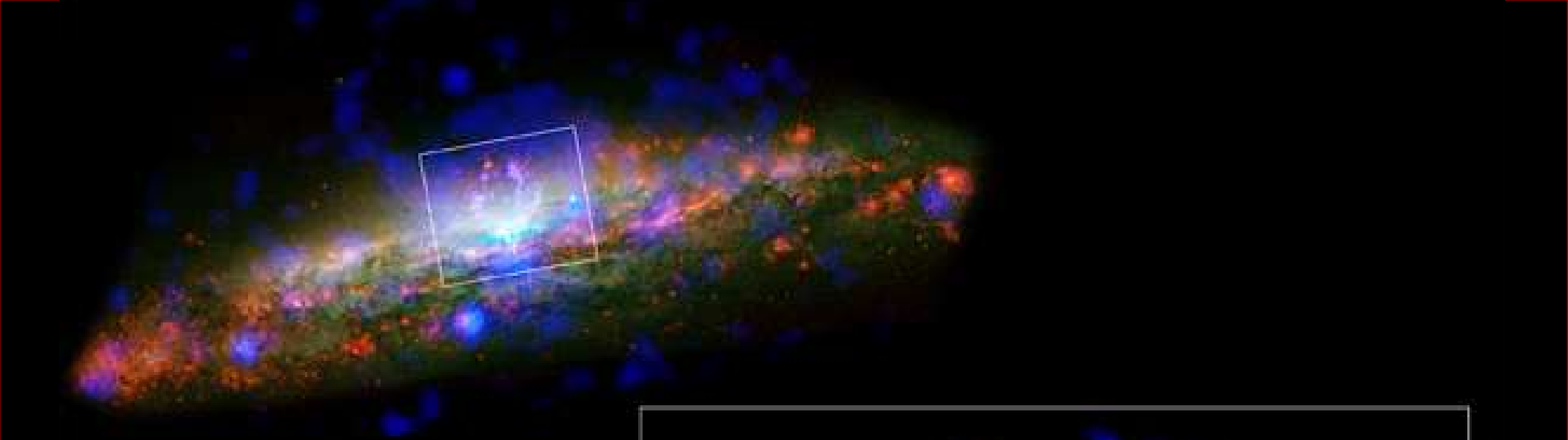


**Highest energies in space above the Sun!
Energy should come from somewhere!
Space is ionized – plasma!**



Where are the highest energies in universe?



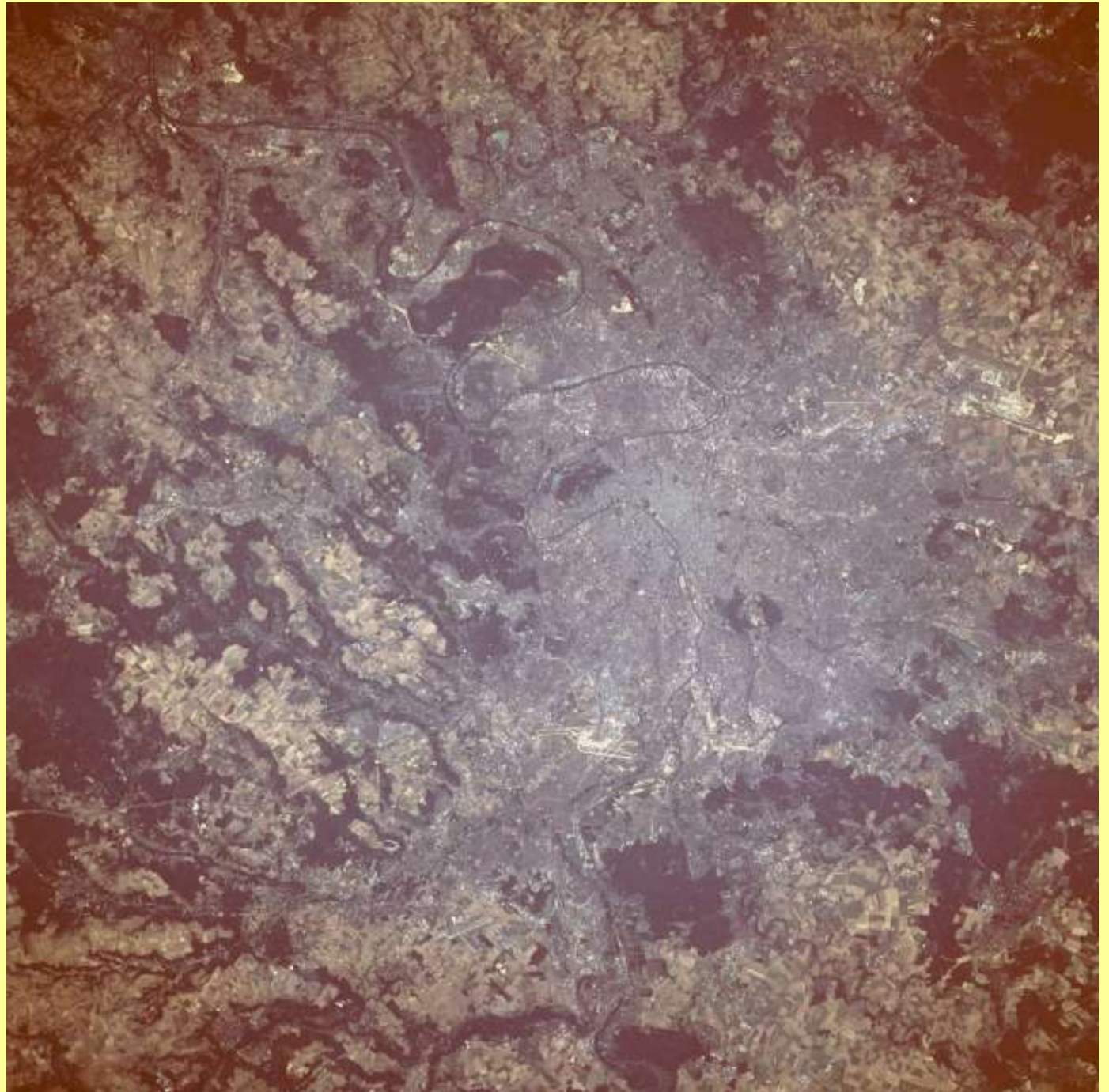
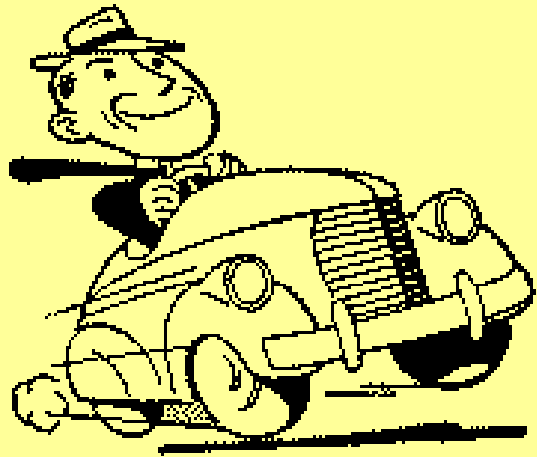


NGC 3079

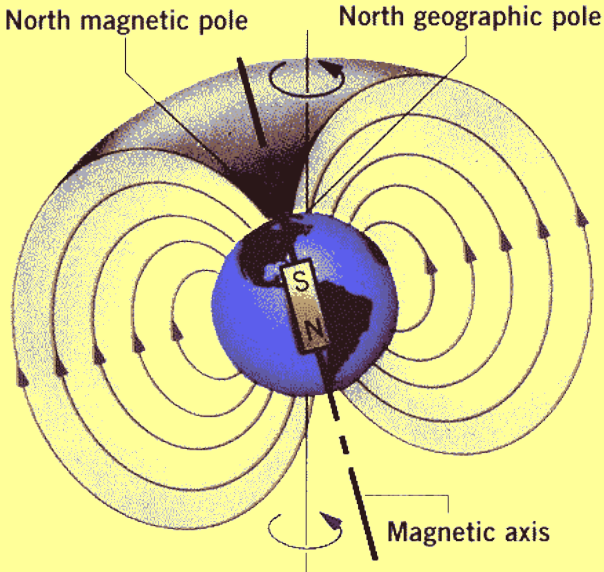
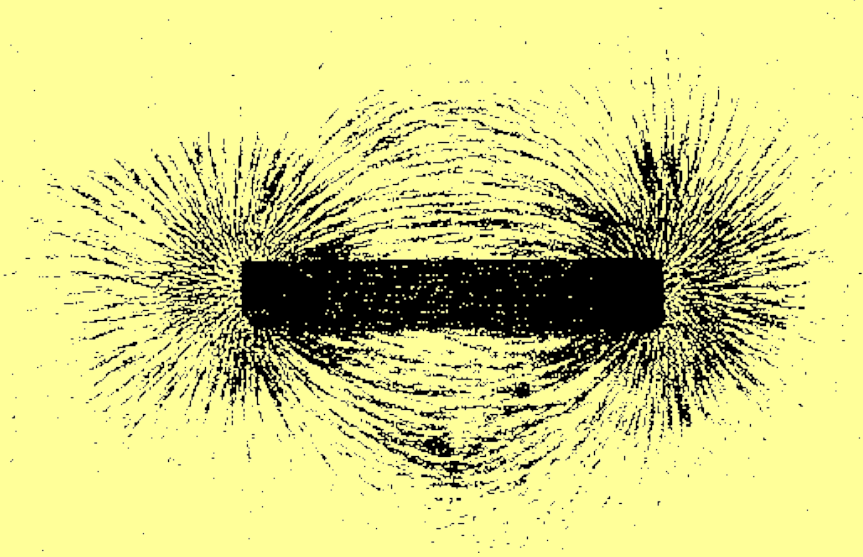
Chandra X ray

Hubble optical

Paris
sunny working day

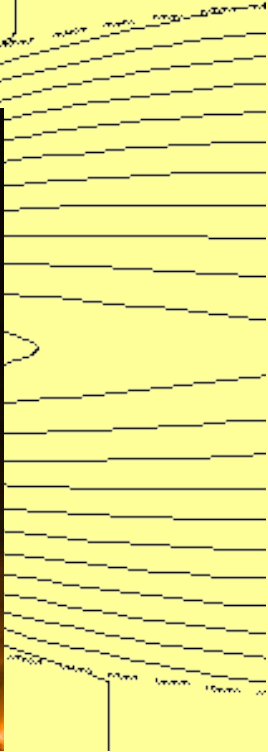
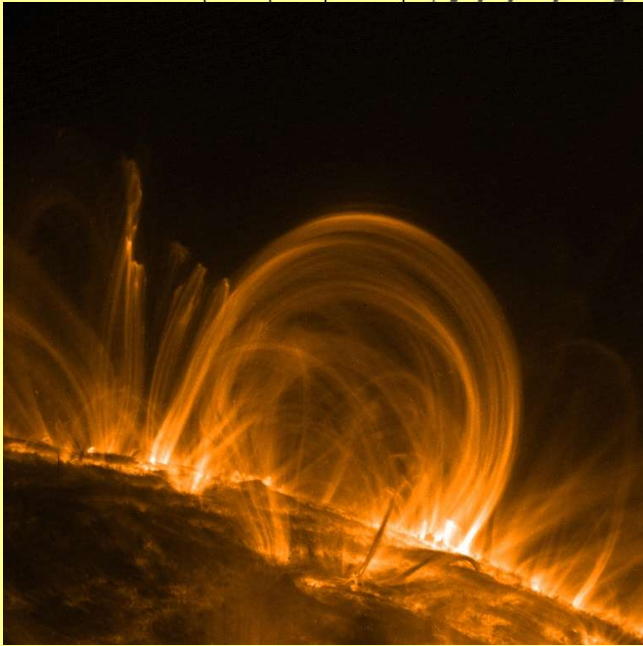
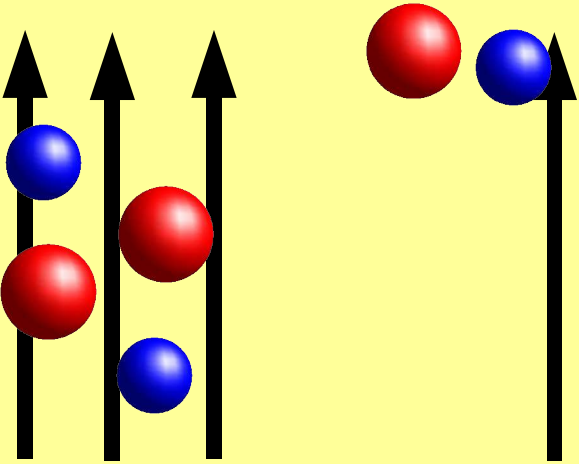


Magnetic field

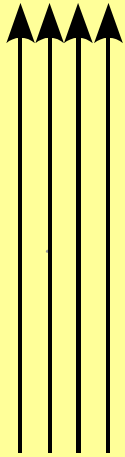


Copyright John Wiley & Sons

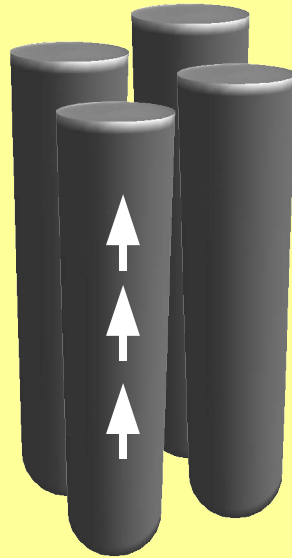
Think in magnetic field lines! .. without ends!



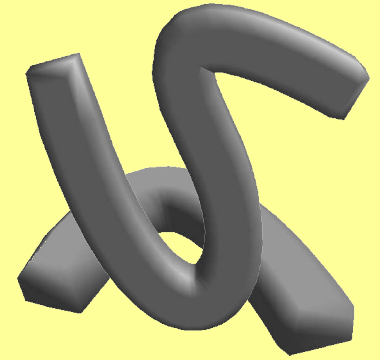
Magnetic flux tube energy W



Magnetic field lines

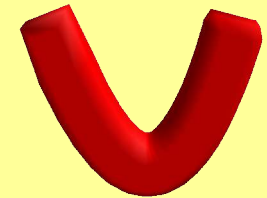
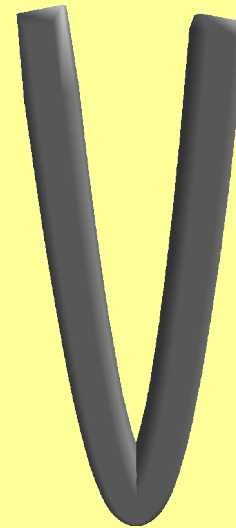


Magnetic flux tubes



Magnetic pressure

$$W > W$$



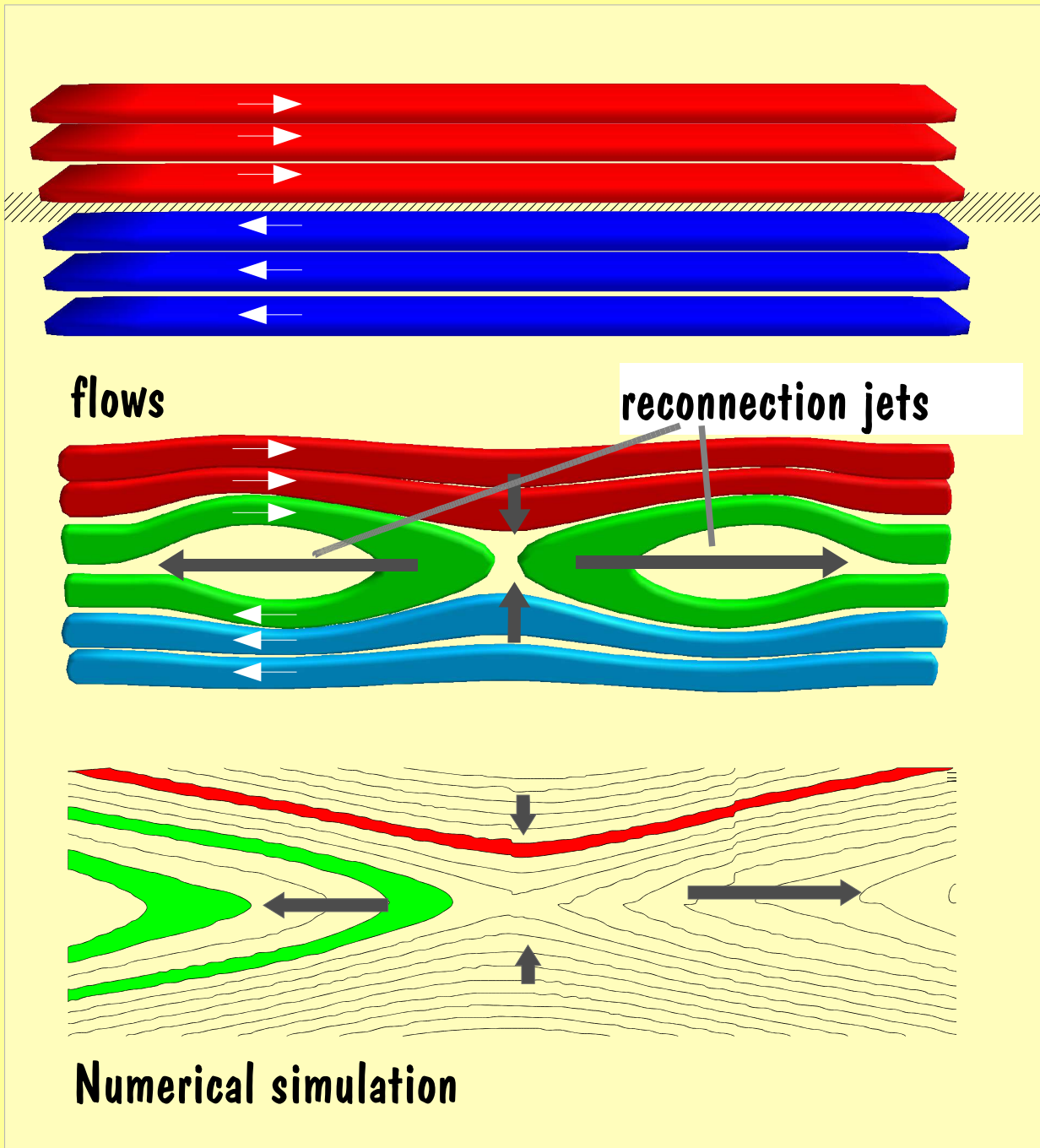
Magnetic tension

Magnetic energy converted into **kinetic energy** of ions and electrons

A deep space photograph showing a vast field of galaxies and stars against a black background. The galaxies are scattered across the frame, some appearing as bright, multi-colored spots, while others are more diffuse and elongated. The stars are small, bright points of light, some with prominent diffraction spikes. The overall scene is a rich, multi-colored tapestry of cosmic objects.

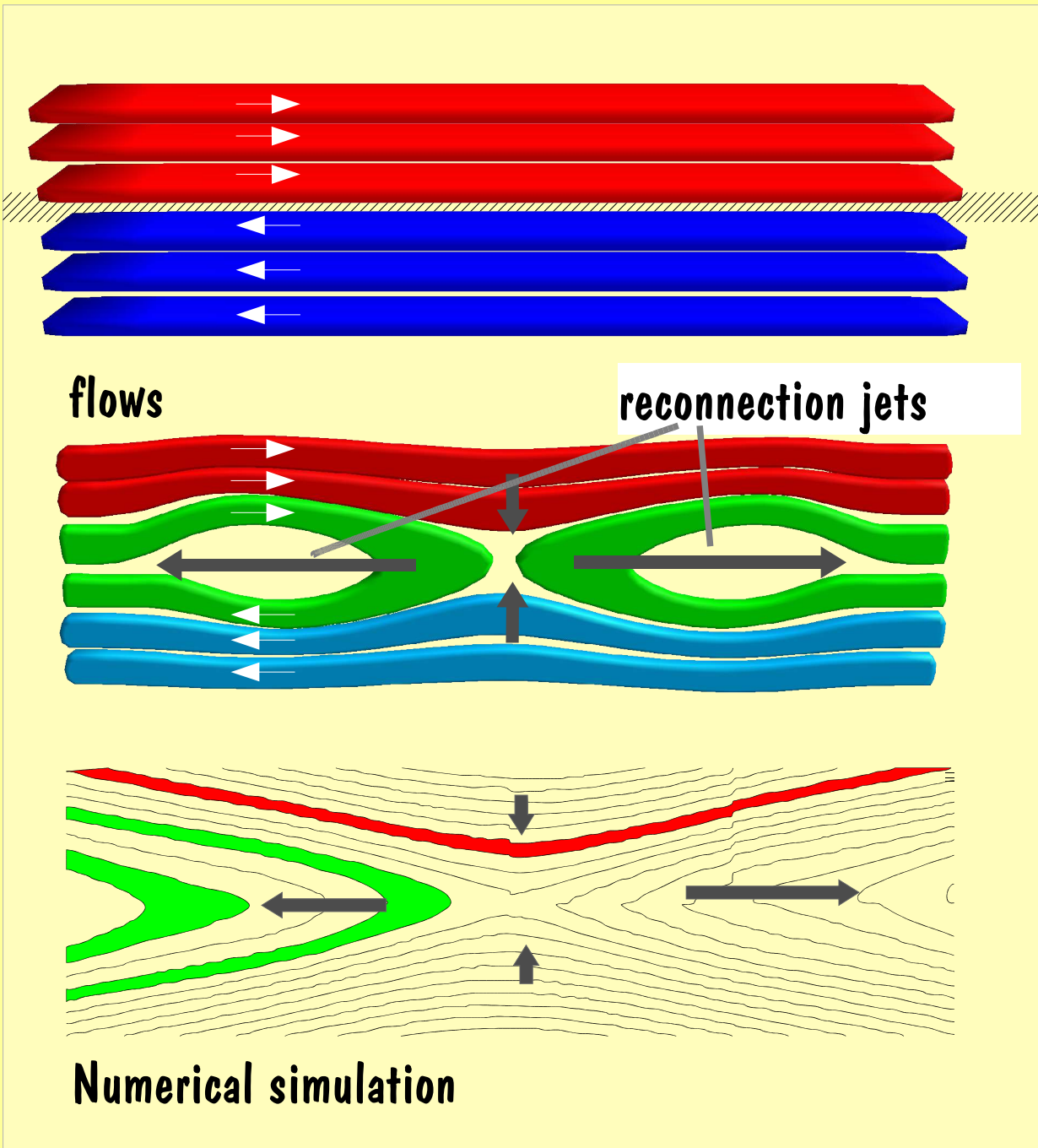
In space there are many places with much more energy in magnetic field than in ions and electrons. **Source of high energy particles?!**

Reconnection – the idea



- ✓ A narrow **current sheet** separates in space regions of different plasmas and oppositely directed magnetic field
- ✓ **Reconnection** – topology changes allowing interconnection of magnetic field lines from both regions. Two ways 1) $E_{\parallel} \neq 0$, 2) $B=0$.
- ✓ **Reconnection** – energy conversion as a result of **reconnection**

Reconnection – the point

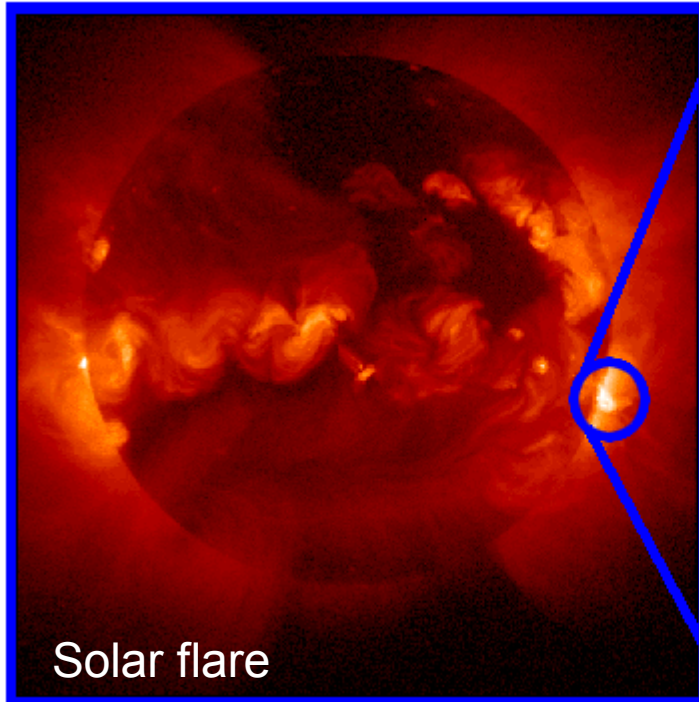


✓ **Topology** – allows energy and mass transport across boundary

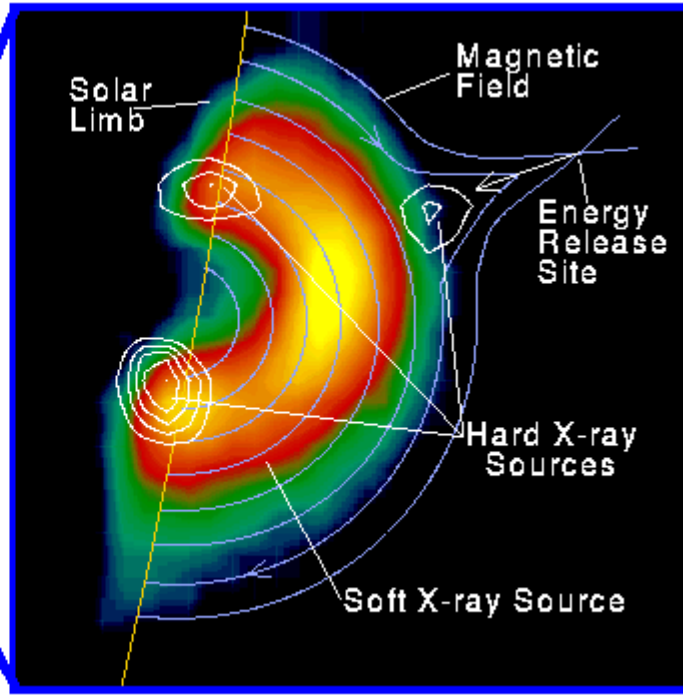
✓ **Energy** – the magnetic field energy is built up by many particles but is converted to the kinetic energy of few which therefore receive **high energies**.

✓ topology vs. energy

Examples from space

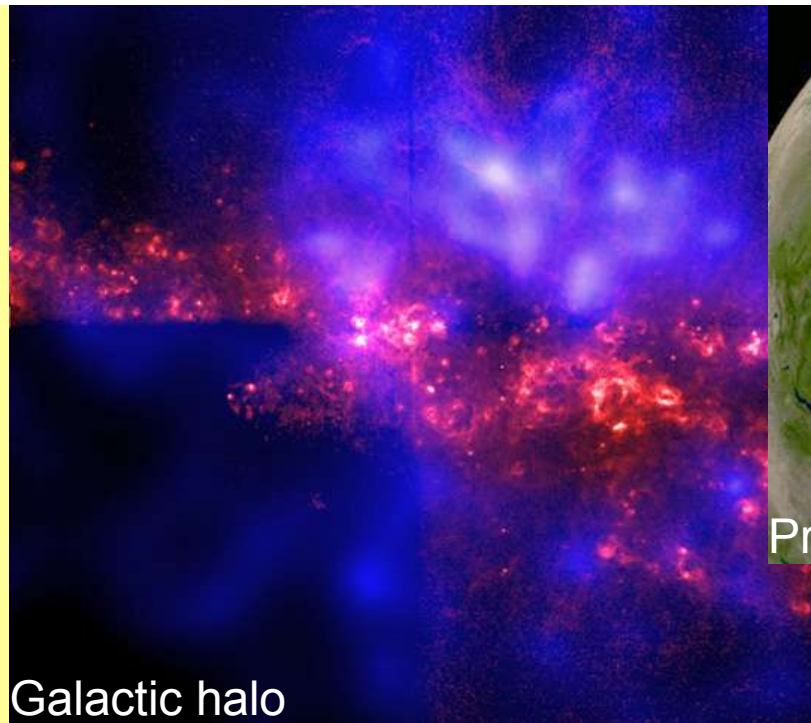


Solar flare

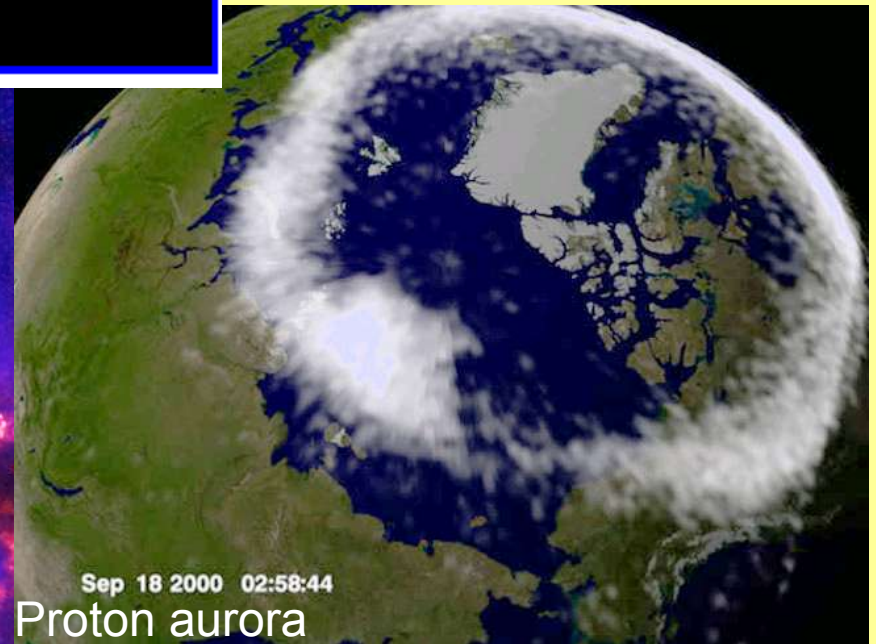


Sun

Galaxies



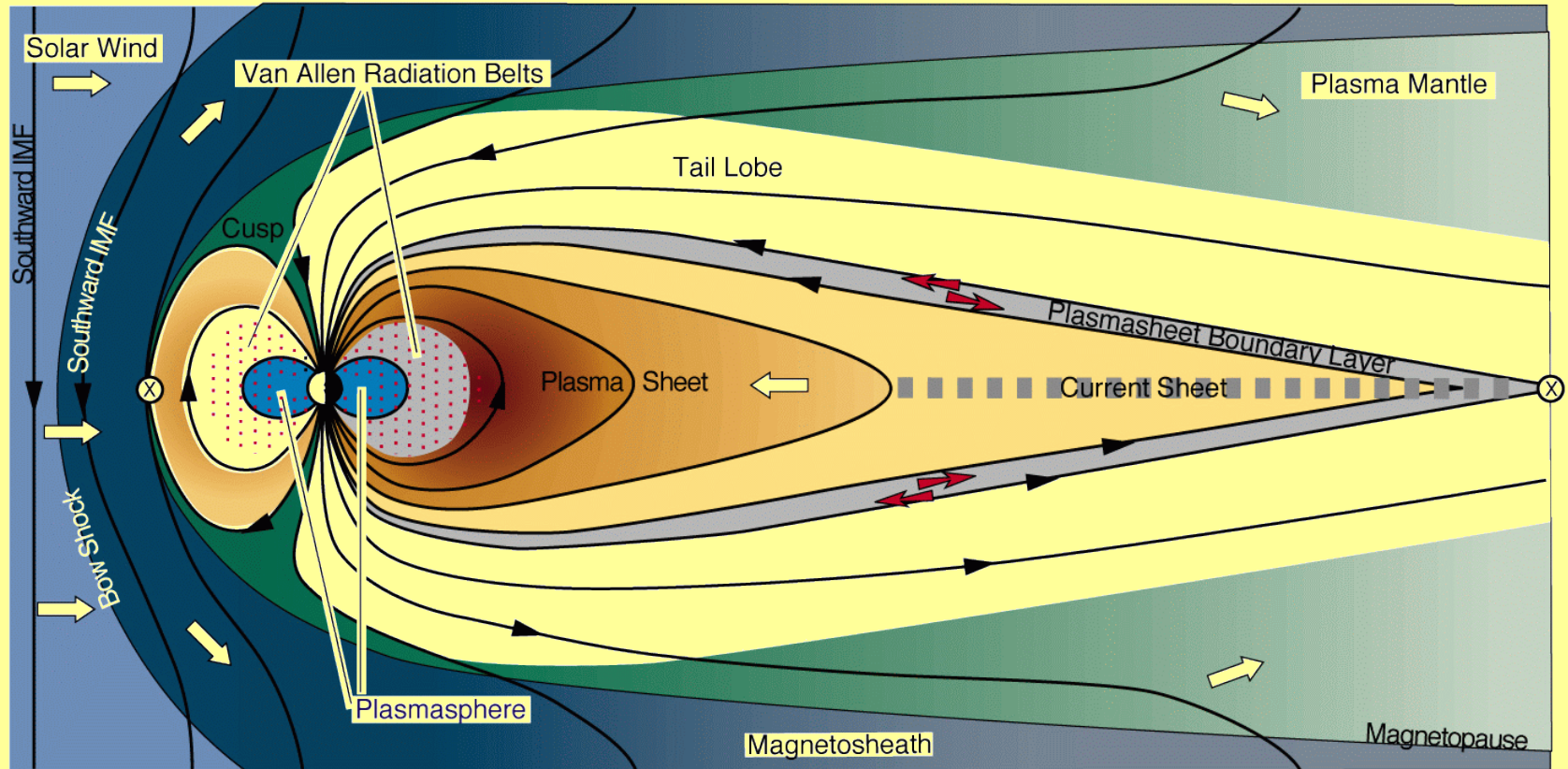
Galactic halo



Sep 18 2000 02:58:44
Proton aurora

Earth

Solar wind and Earth magnetosphere

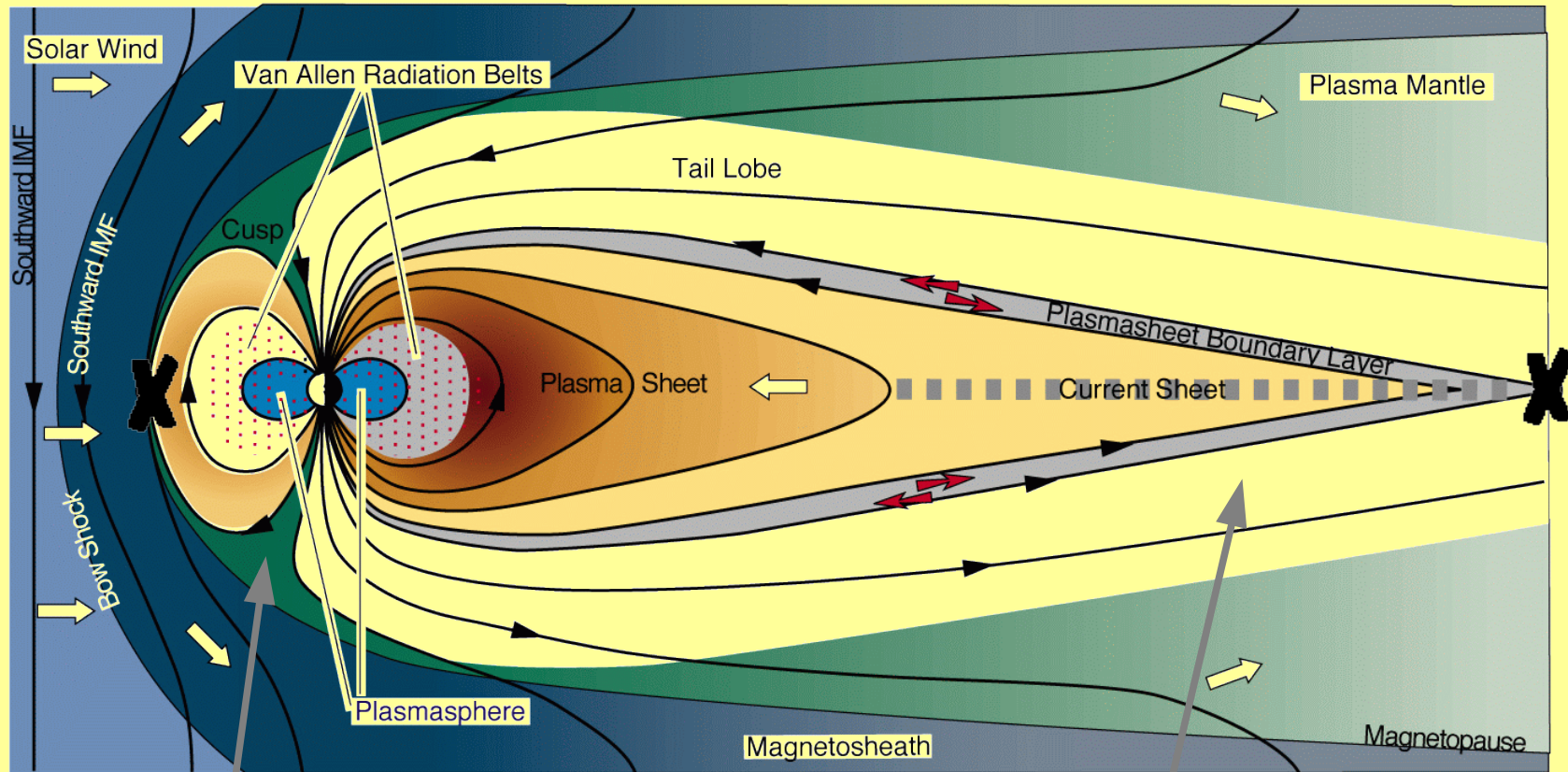


Solar wind and Earth magnetosphere

Reconnection

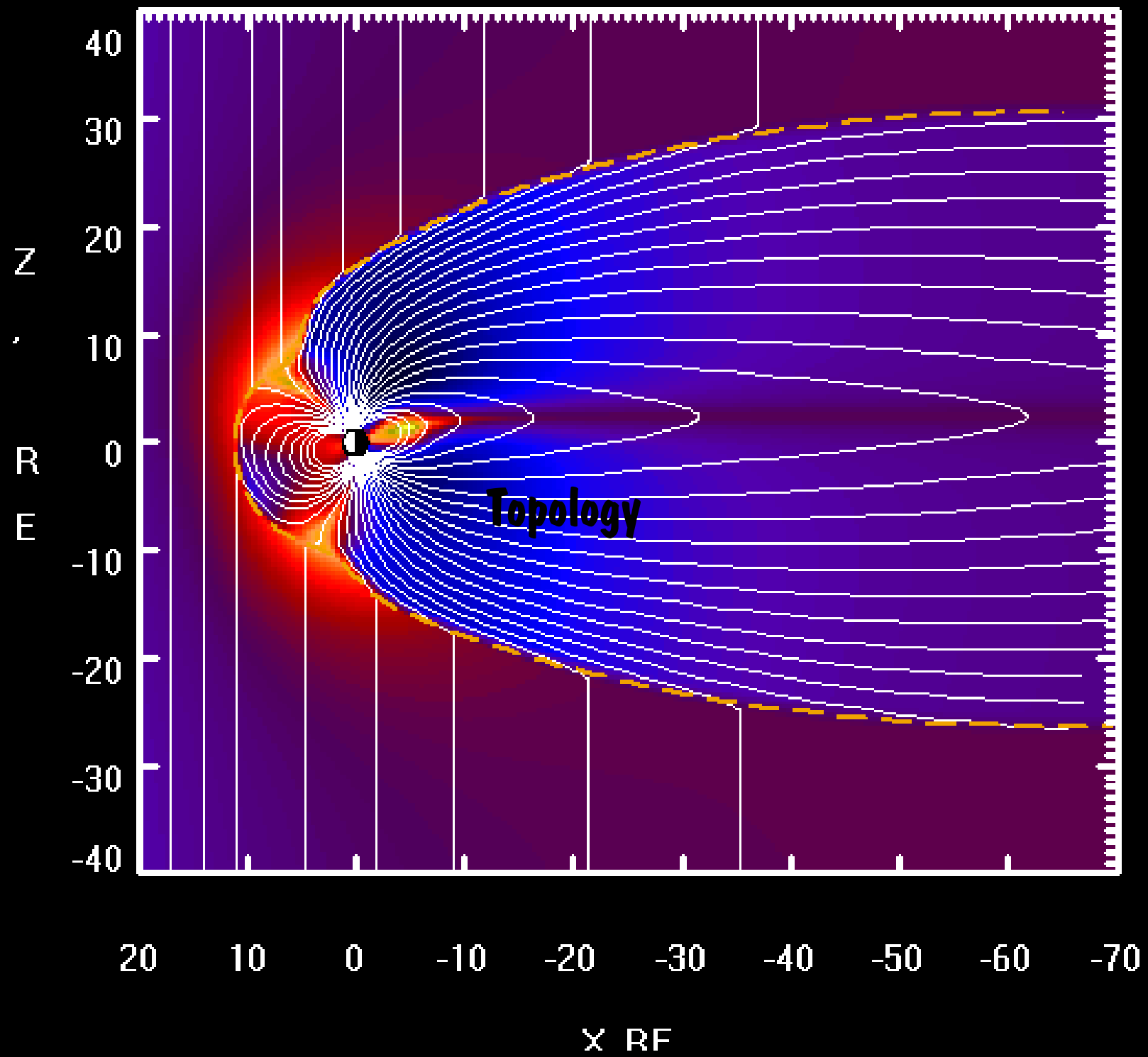
Topology

Energy



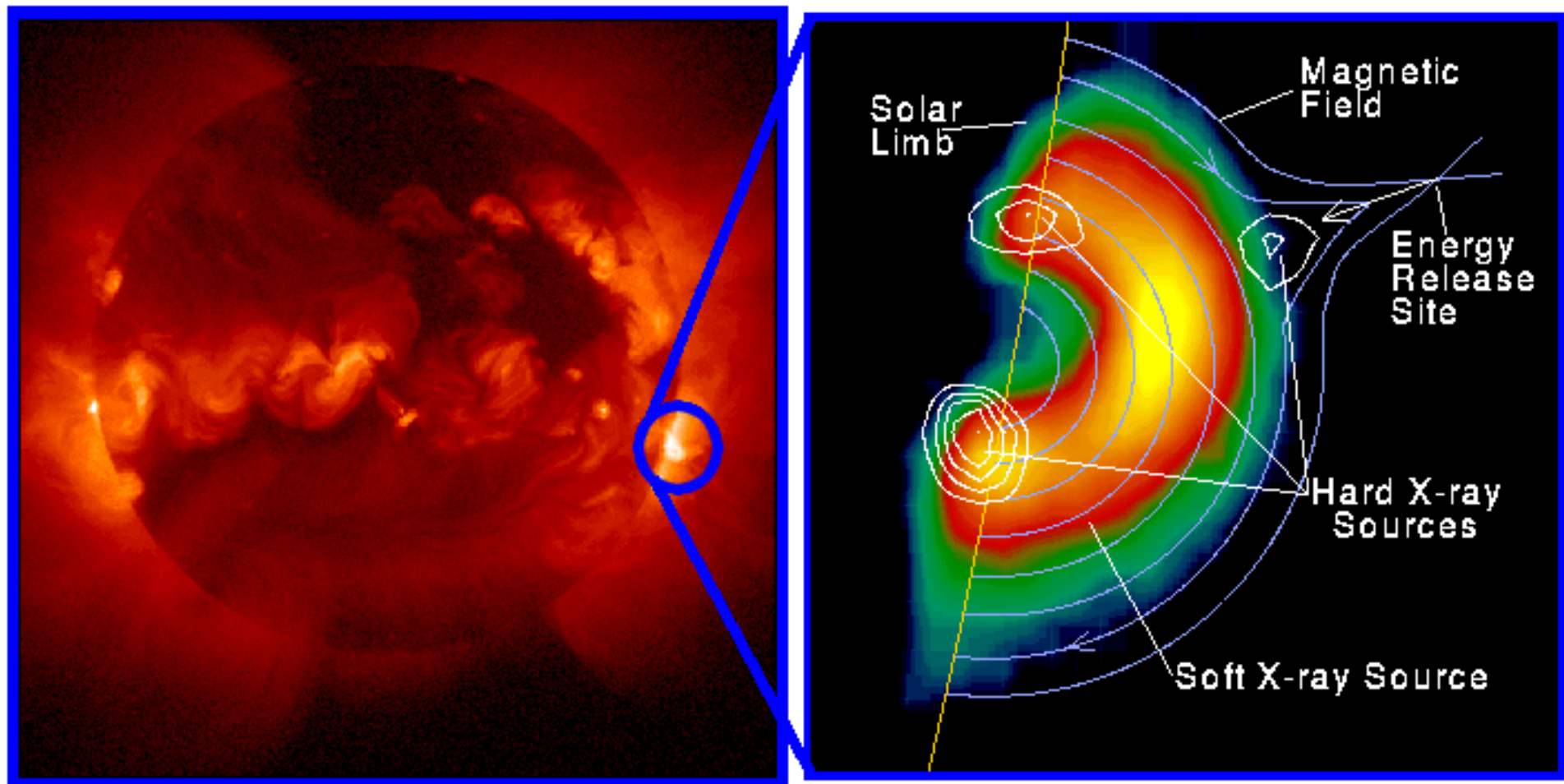
Magnetopause
energy transfer

Magnetotail
energy storage (B)



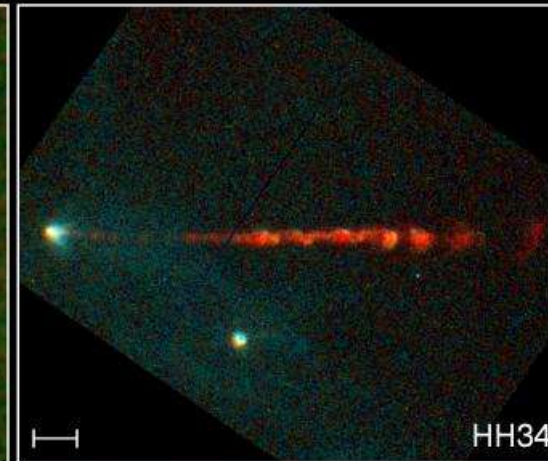
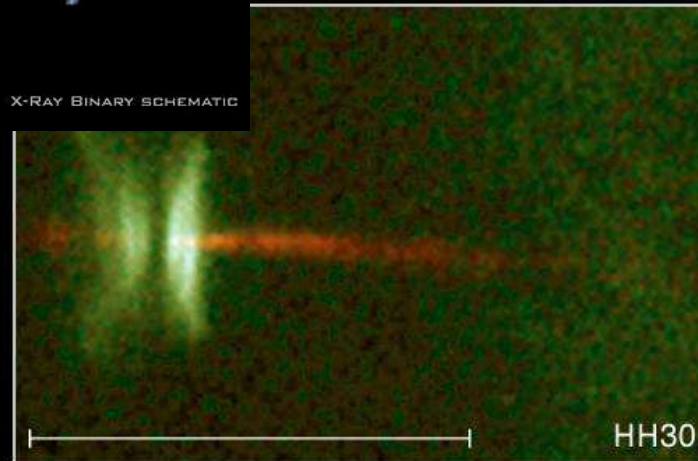
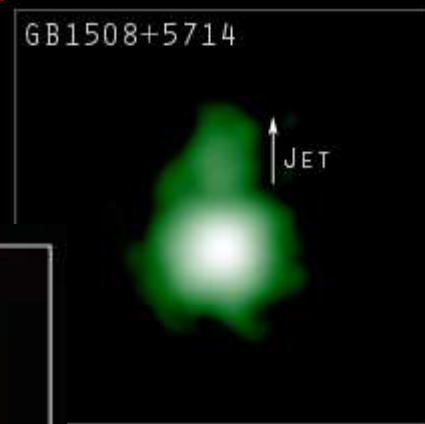
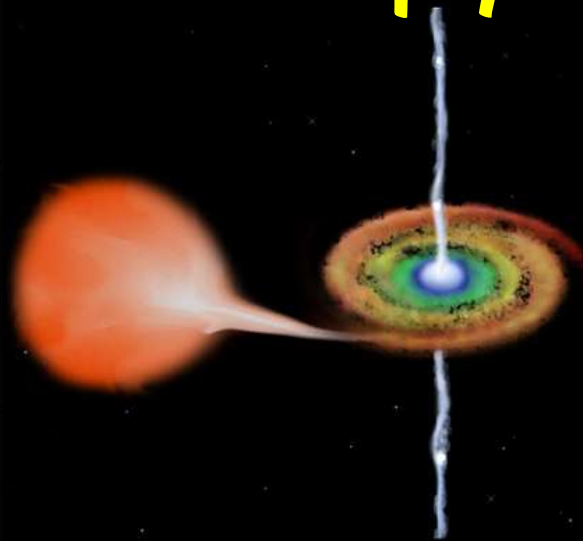
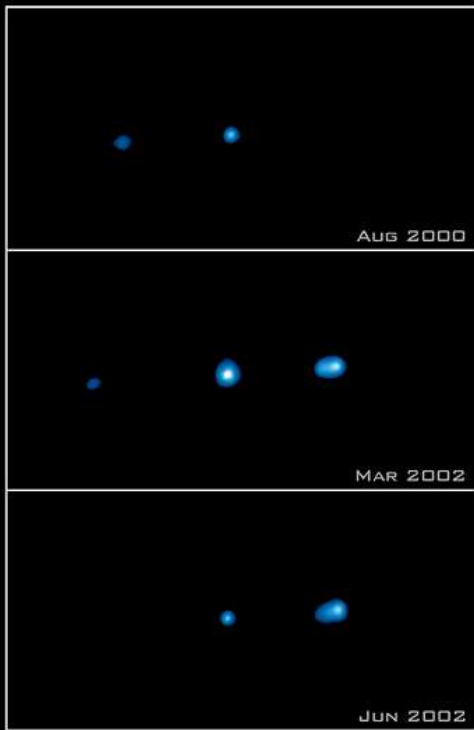
Solar flares

Xflare movie



Yohkoh X-ray Image of a Solar Flare, Combined Image in Soft X-rays (left) and Soft X-rays with Hard X-ray Contours (right). Jan 13, 1992.

Astrophysical jets



Jets from Young Stars

PRC95-24a · ST Scl OPO · June 6, 1995

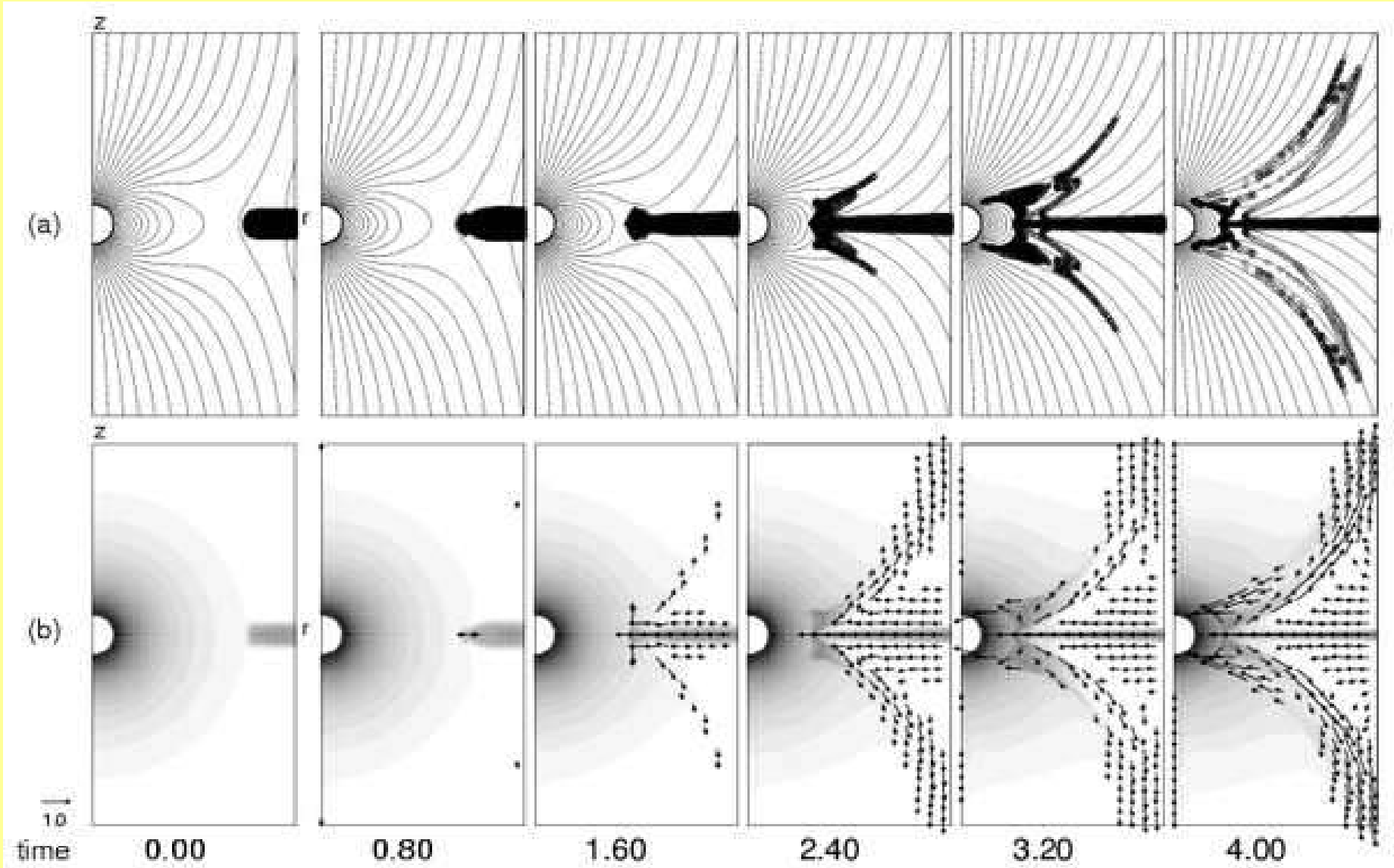
C. Burrows (ST Scl), J. Hester (AZ State U.), J. Morse (ST Scl), NASA

HST · WFPC2

Optical jets from young stars

Numerical simulations

[Hirose et al. 1997]



On a way to understand reconnection

✓ Theory

- do the right approximations
- give up and turn to numerical simulations

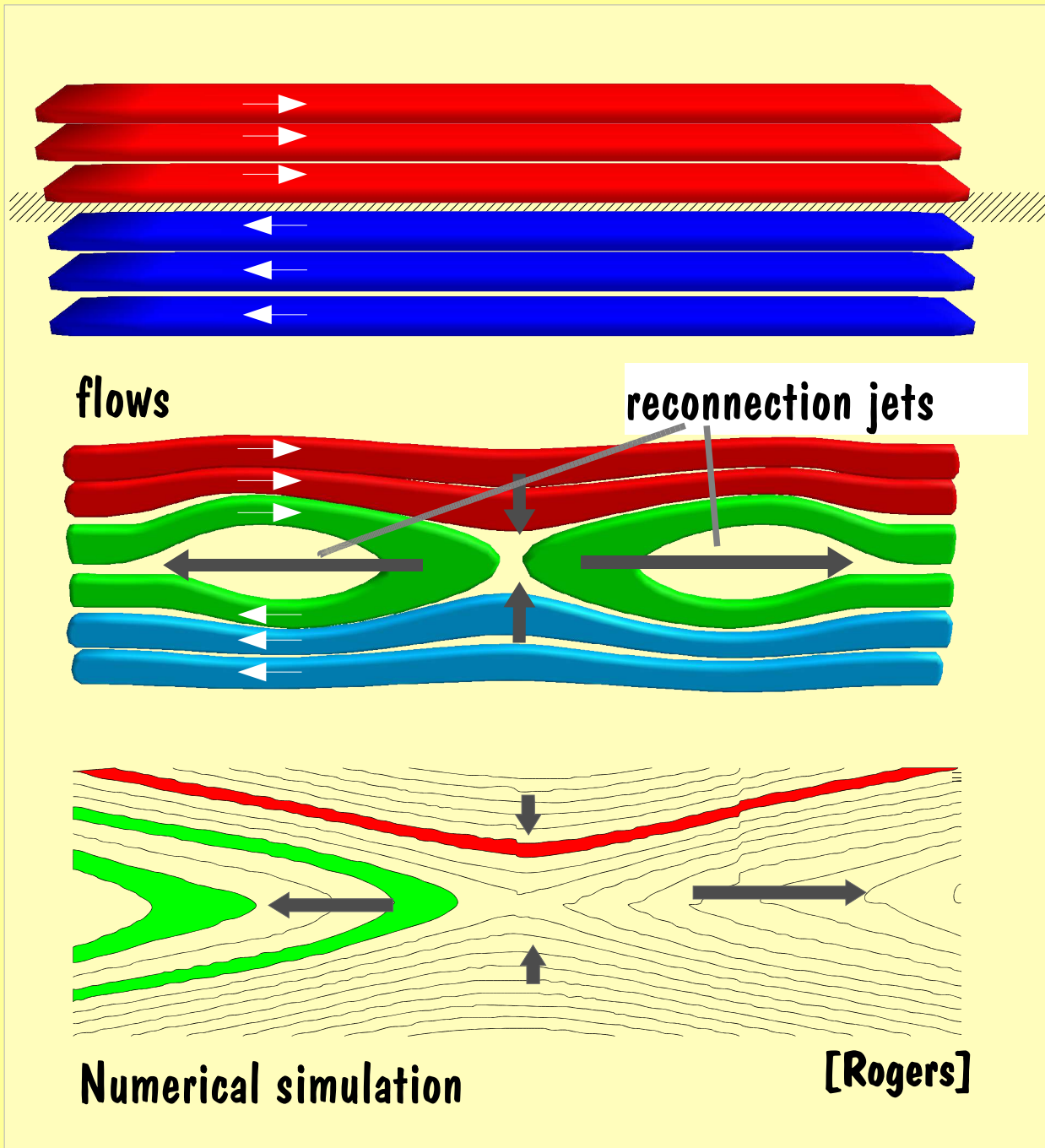
✓ Numerical simulations

- take the right equations (MHD, HMHD, particles)
- put into the code and study results

✓ Experiment

- space in situ / astro / laboratory
- **most detailed observations from space in situ!**

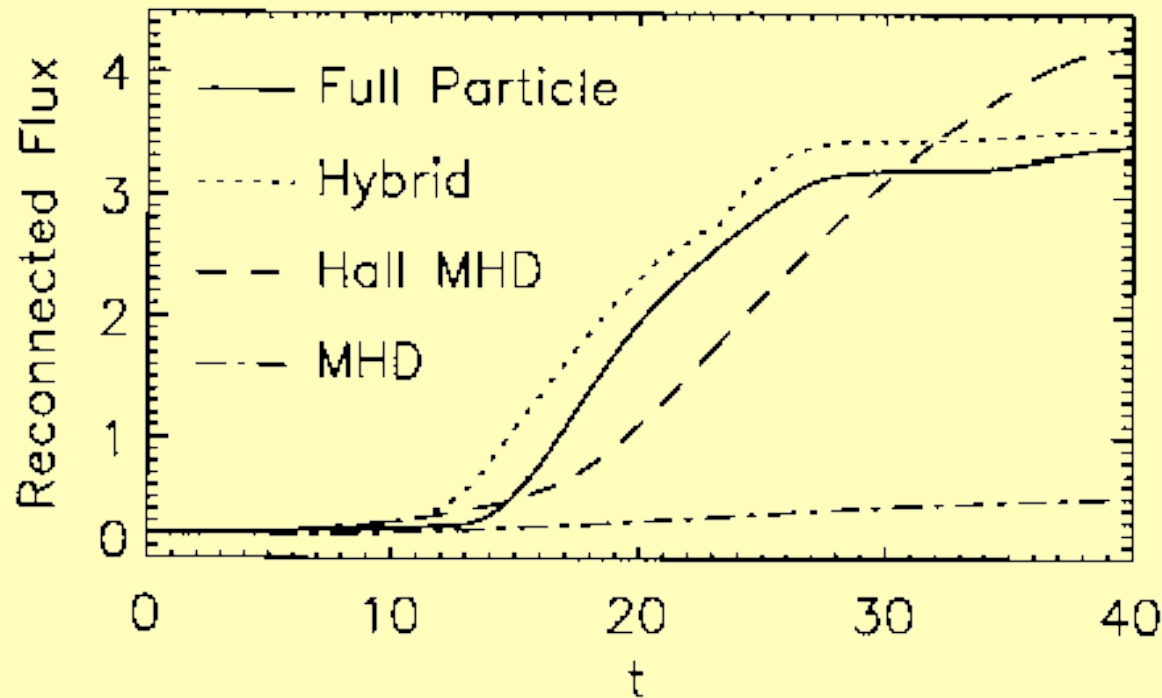
Reconnection – topics



- ✓ Structure
- ✓ Continuity
- ✓ Remote sensing
- ✓ Reconnection rate
- ✓ Guide field
- ✓ Relative importance
- ✓ Acceleration mechanisms (E)
- ✓ X-line

Numerical simulations

[Birn et al., 2001]

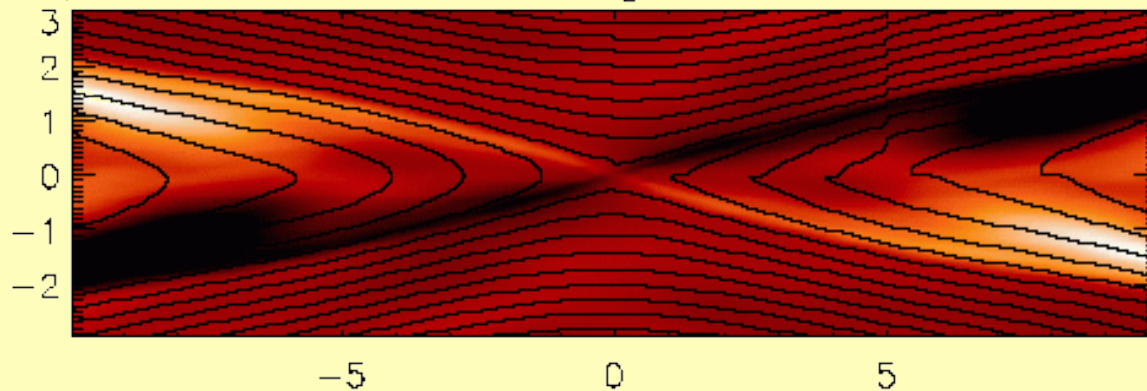


Reconnected flux
GEM challenge

Hall term important!

- whistlers
- Kinetic Alfvén waves

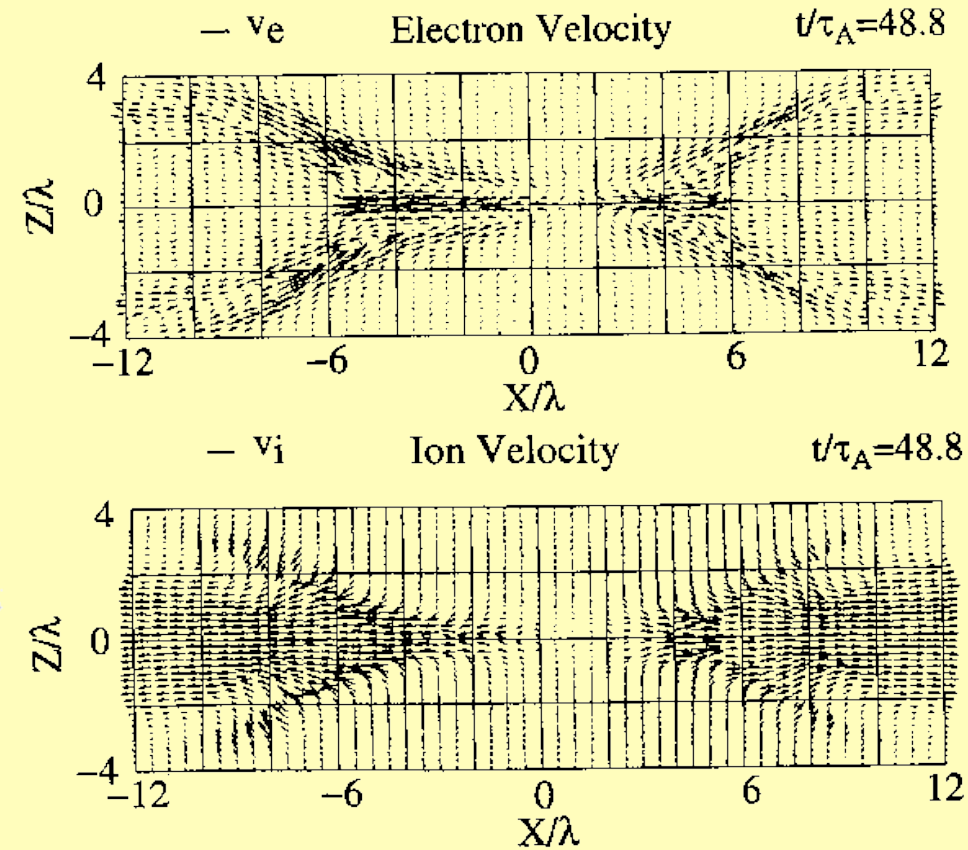
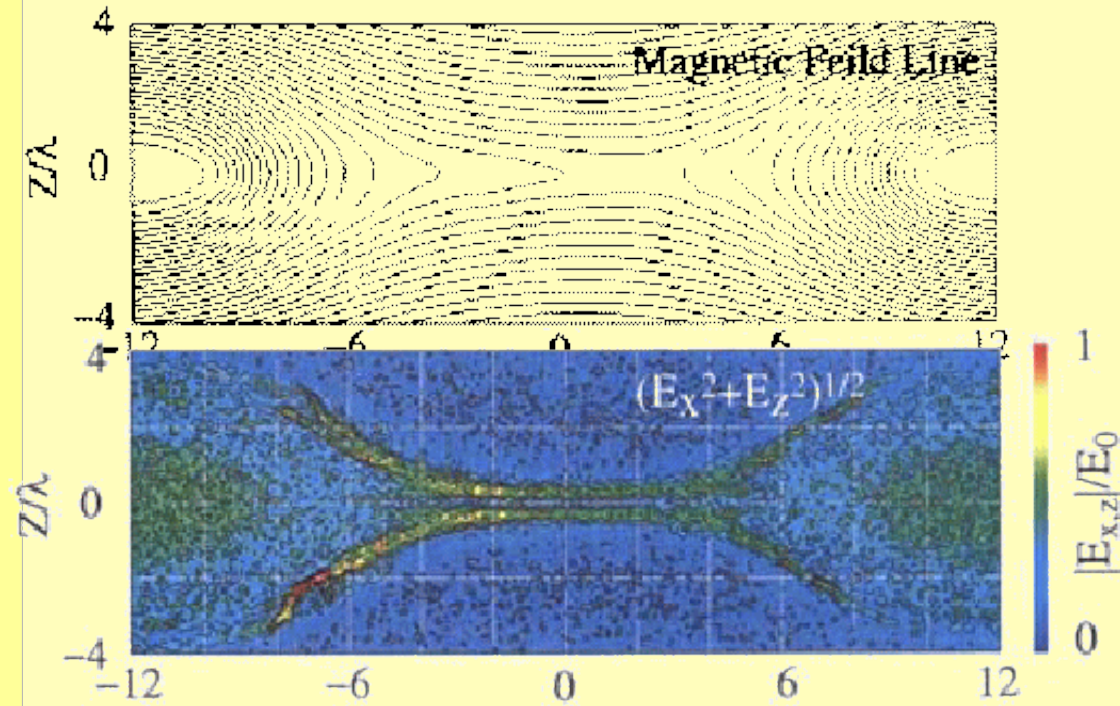
[Rogers]



out of plane magnetic field
- quadrupolar structure

Numerical simulations

[Hoshino et al., 2001]

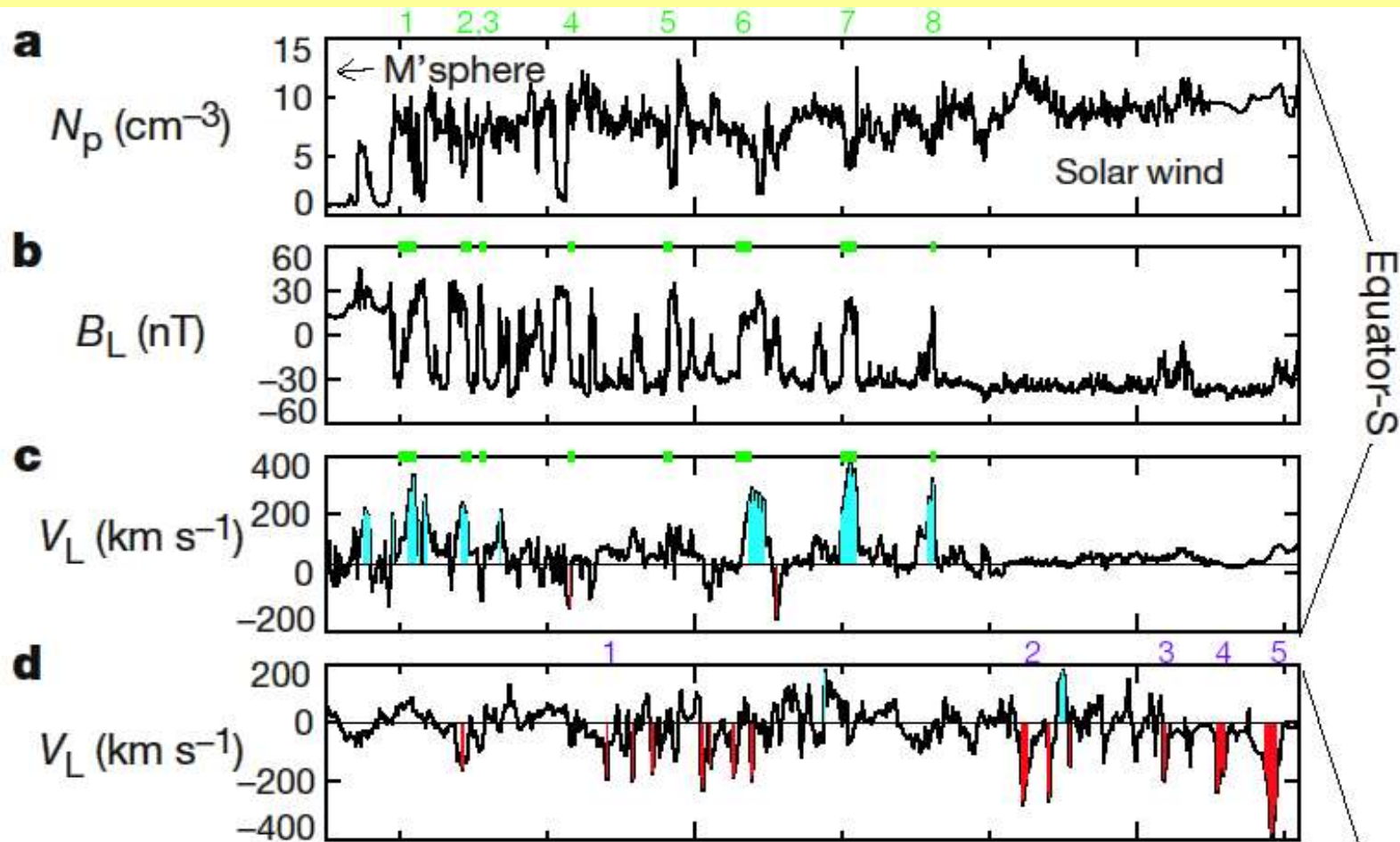


Separatrices

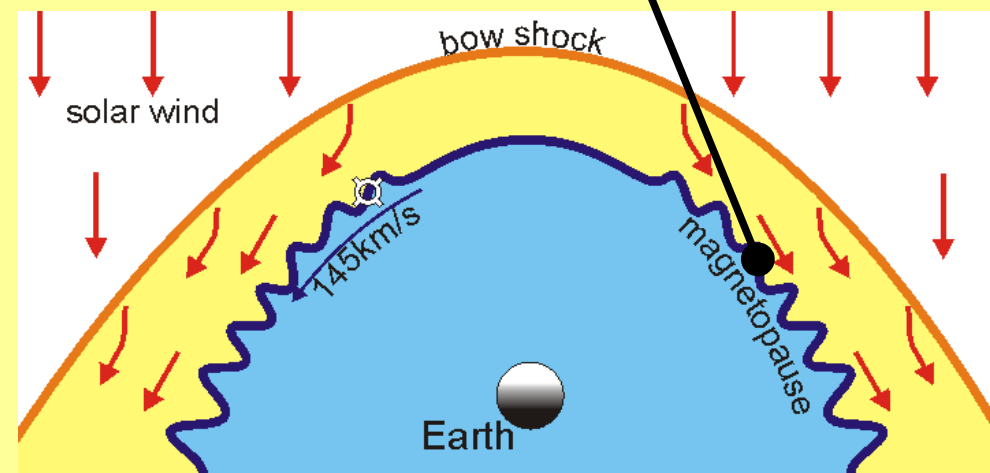
- strong electric fields
- strong parallel currents
- electron beams

Reconnection & space physics

[Phan et al., 2000]

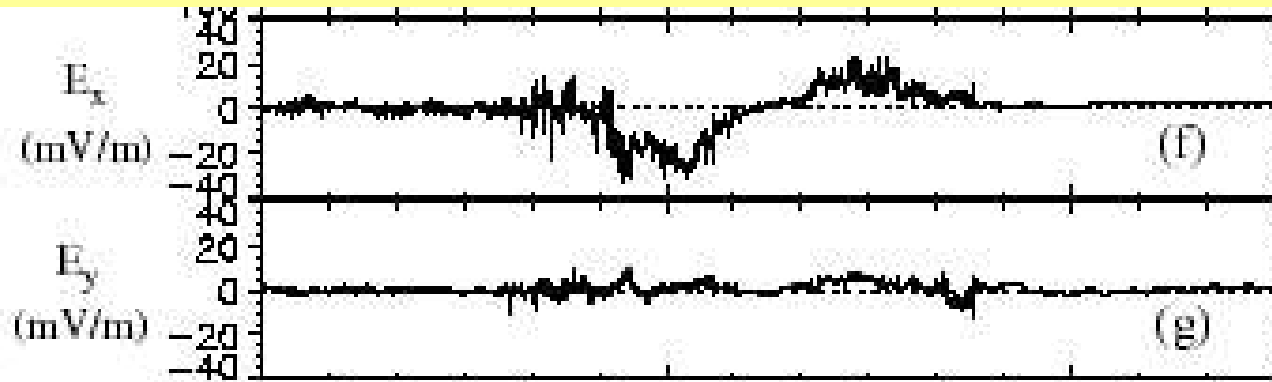
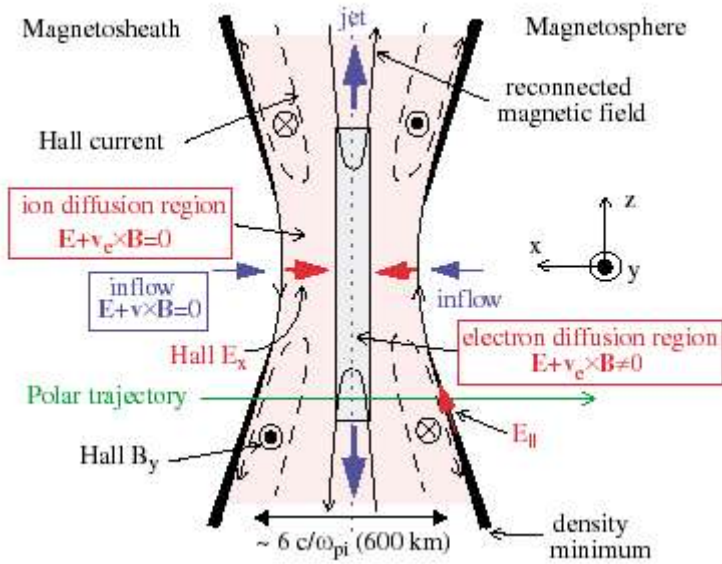


- ✓ Reconnection jets
- ✓ Continuity



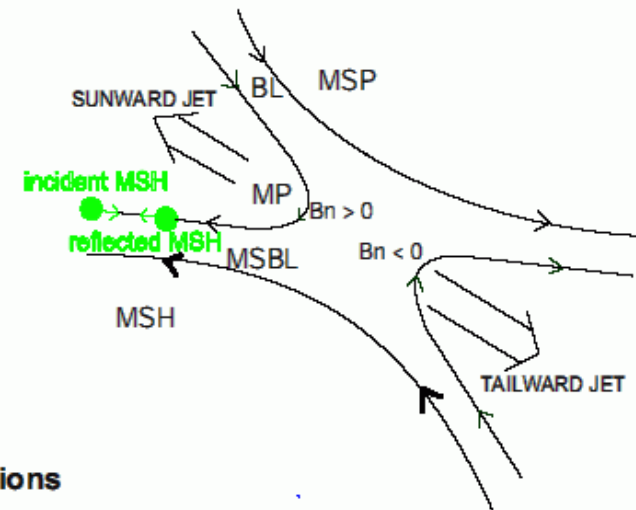
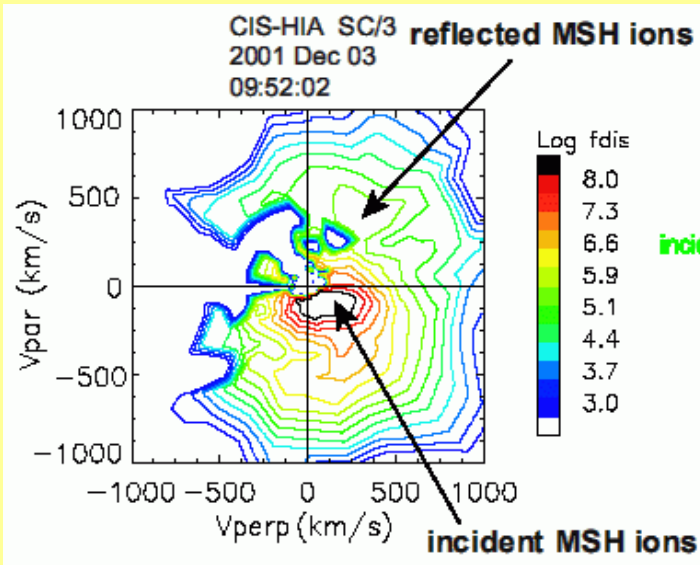
Reconnection & space physics

Why space in situ : we can measure local magnetic and electric fields, as well as ion and electron distribution functions.

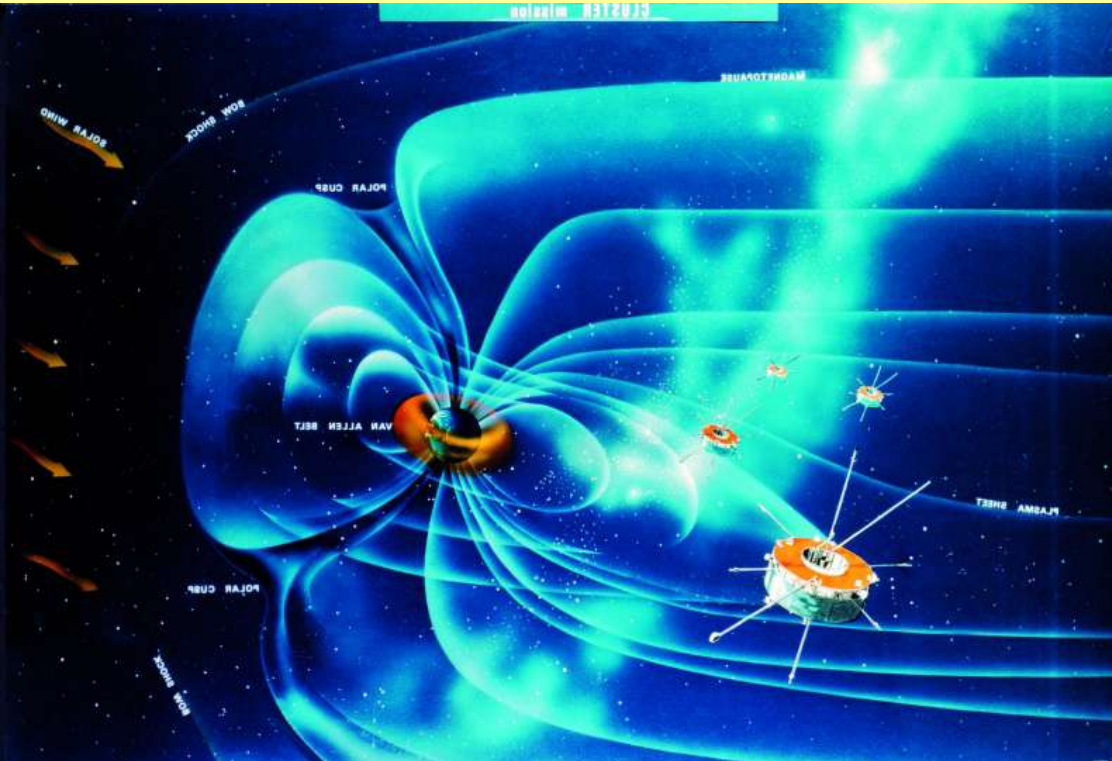


[Mozer et al., 2002]

[Retino et al., 2004]

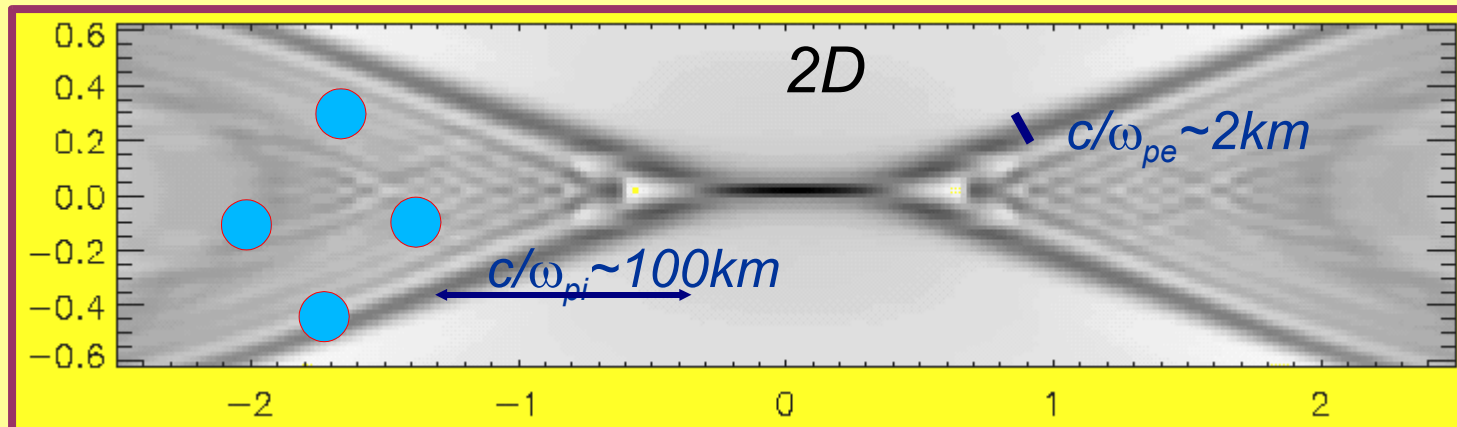


Cluster chasing X-line



- ✓ 4 s/c ESA "Cornerstone"
- ✓ 2000-2005
- ✓ First time we can reliably solve time-space ambiguity
- ✓ First time reliable measurements of current

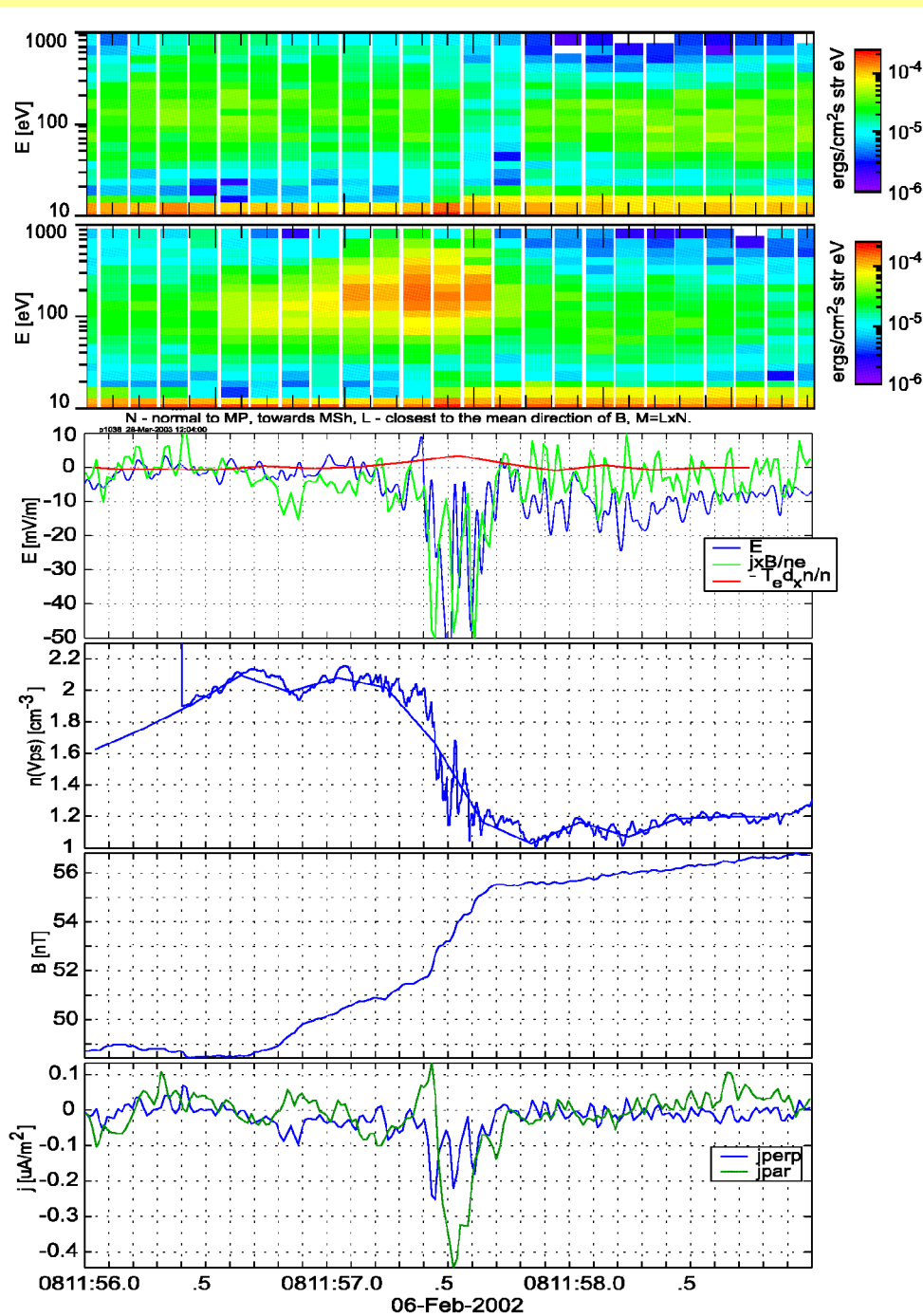
Current



[Rogers et al., 2000]

Cluster, separatrixes far from X-line

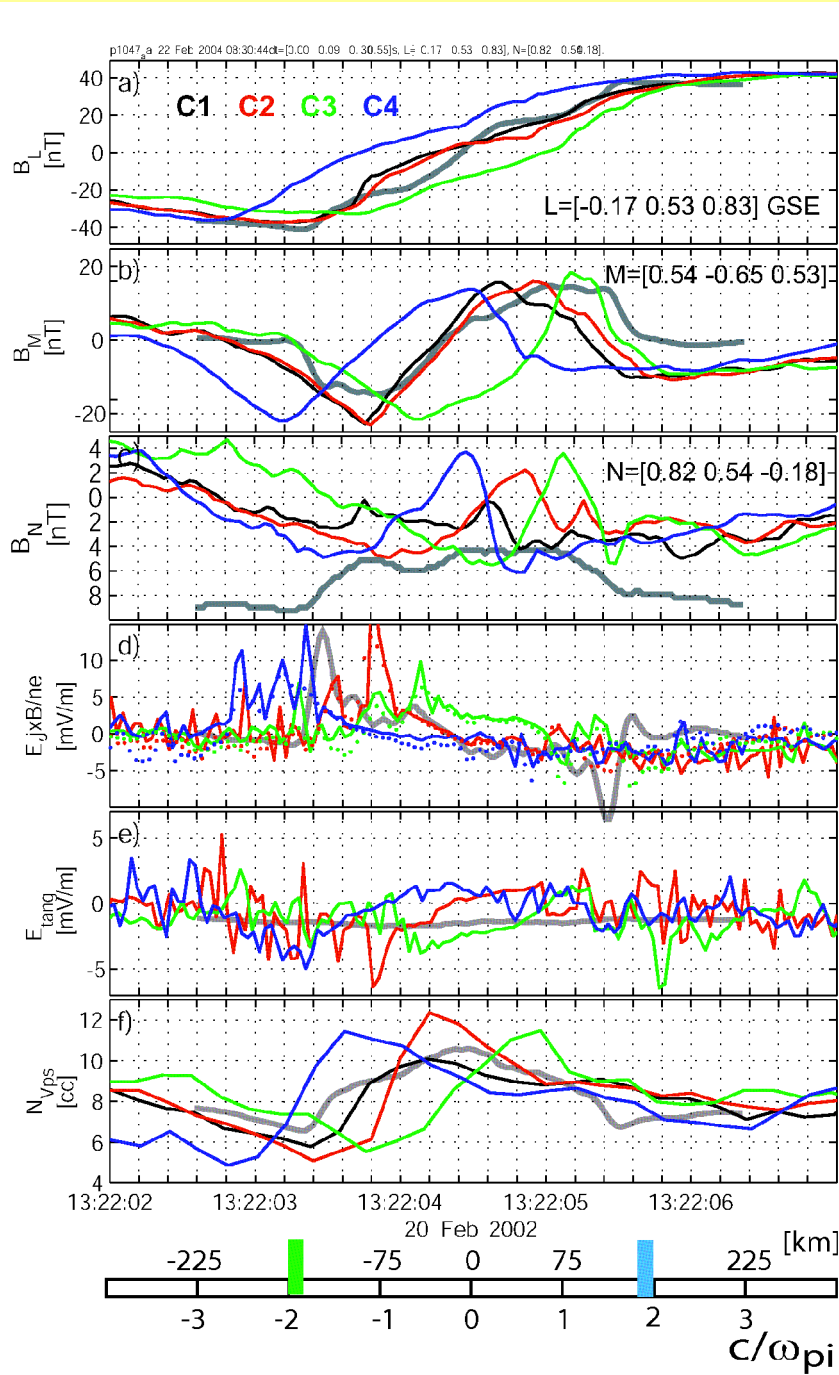
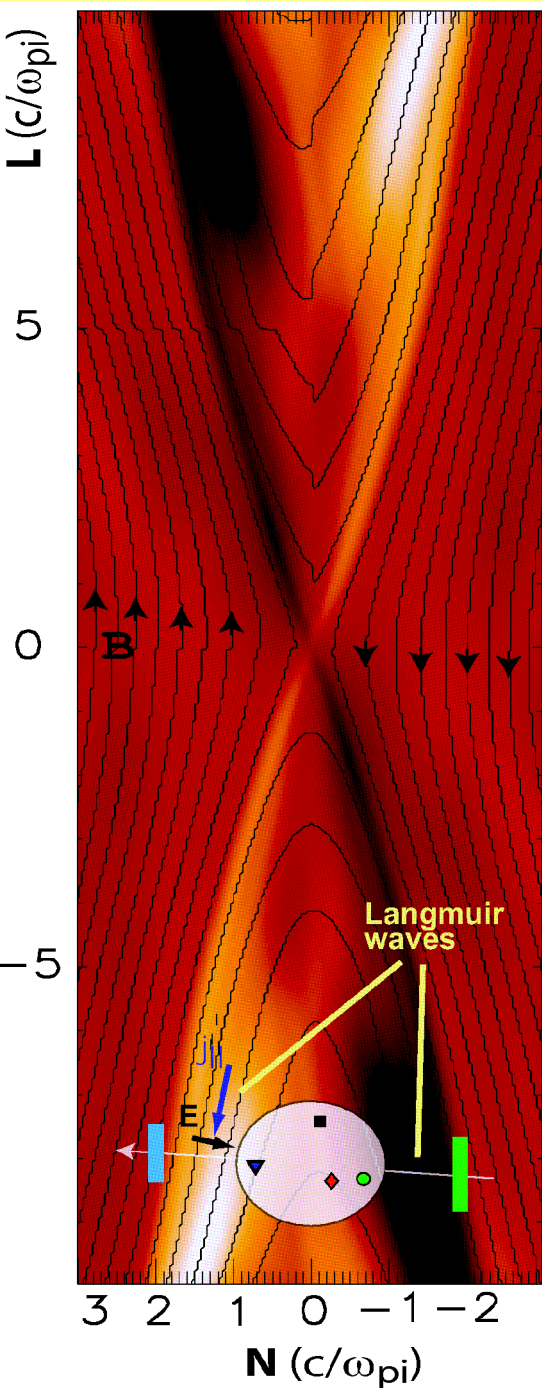
[André et al., 2004]



- ✓ Narrow strong current sheet
- ✓ e- beams
- ✓ Hall dynamics

Cluster, close to X-line

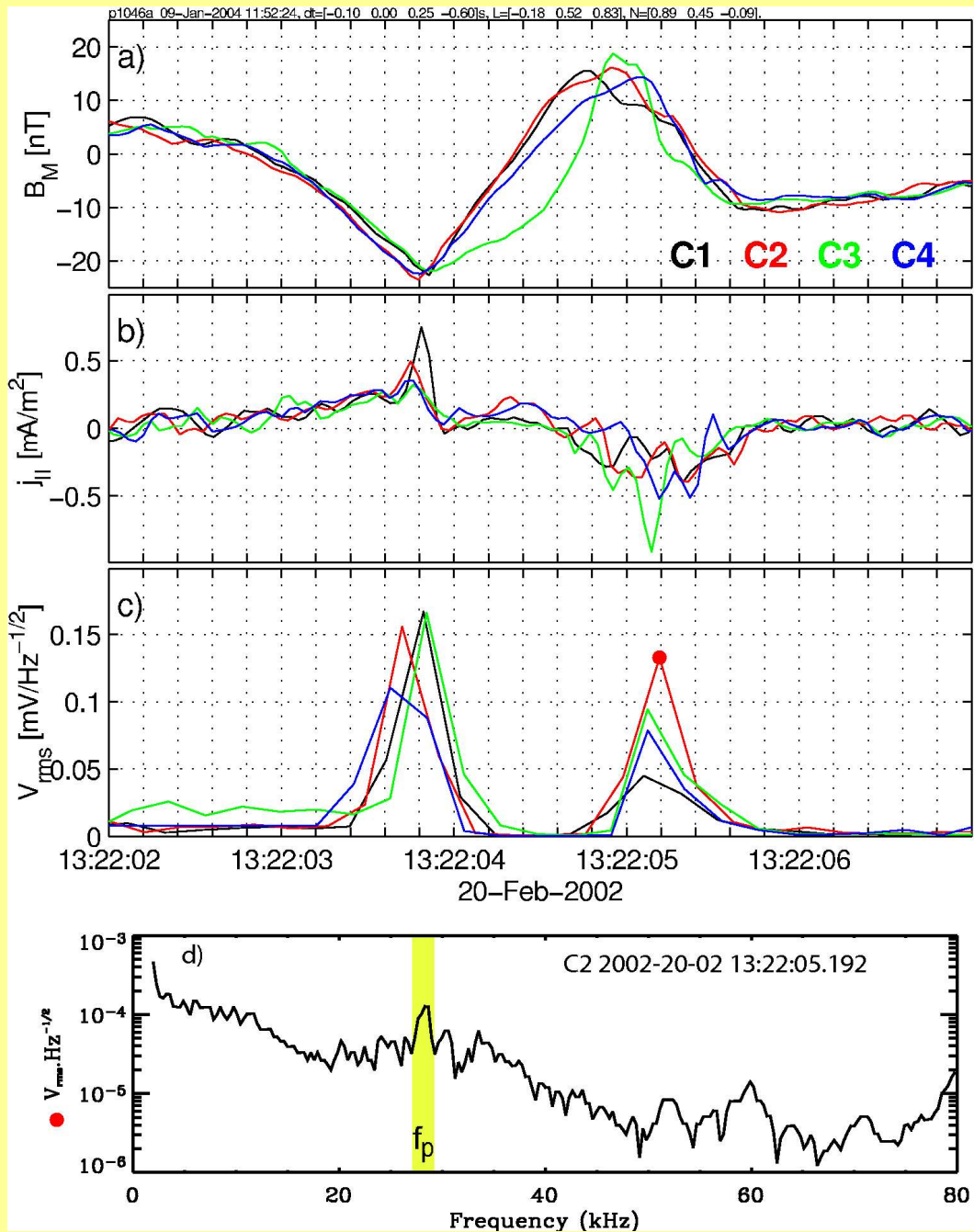
[Vaivads et al., 2004]



- ✓ Reliable structure estimate
- ✓ Hall dynamics
- ✓ Fast reconnection

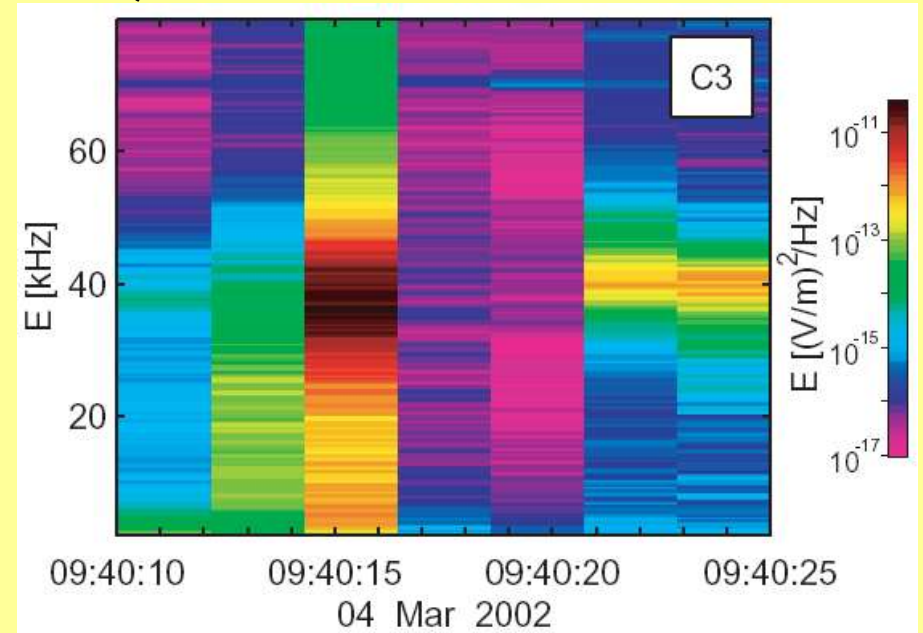
Separatrices, j_{\parallel} and HF waves

[Vaivads et al., 2004]



- ✓ Strong parallel currents
- ✓ Langmuir/upper hybrid emissions
- ✓ The role of waves?
(anomalous resistivity, diffusion)

[Khotyaintsev et al., 2004]



Reconnection summary:

- ✓ Magnetic field **topology** changes and **energy** conversion
- ✓ Reconnection **jets**, plasma **energization**
- ✓ PLASMA (low energy) $> B >$ plasma (**HIGH ENERGY**)
- ✓ In space around planets, stars, black holes, ...
- ✓ We start to understand **details**